

Eventually once true ground state ψ_0 is reached, then $\hat{a}_- \psi_0$ must be 0. We can use this condition to find ψ_0 .

$$\underbrace{\frac{1}{\sqrt{2\hbar m\omega}} \left(\hbar \frac{d}{dx} + m\omega x \right)}_{\hat{a}_-} \psi_0 = 0 \quad \longrightarrow \quad \frac{d\psi_0}{dx} = -\frac{m\omega}{\hbar} x \psi_0$$

$$\int \frac{d\psi_0}{\psi_0} = -\frac{m\omega}{\hbar} \int x dx \quad \Rightarrow \quad \ln \psi_0 = -\frac{m\omega}{2\hbar} x^2 + \text{constant},$$

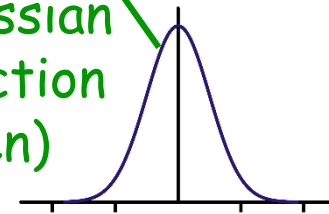
$$\psi_0(x) = A e^{-\frac{m\omega}{2\hbar} x^2}.$$

After normalization:
(left as exercise, use
Gaussian integrals in back
cover of book)

$$\psi_0(x) = \left(\frac{m\omega}{\pi \hbar} \right)^{1/4} e^{-\frac{m\omega}{2\hbar} x^2}$$

Check!

Gaussian
function
(even)



What is the energy E_0 ?

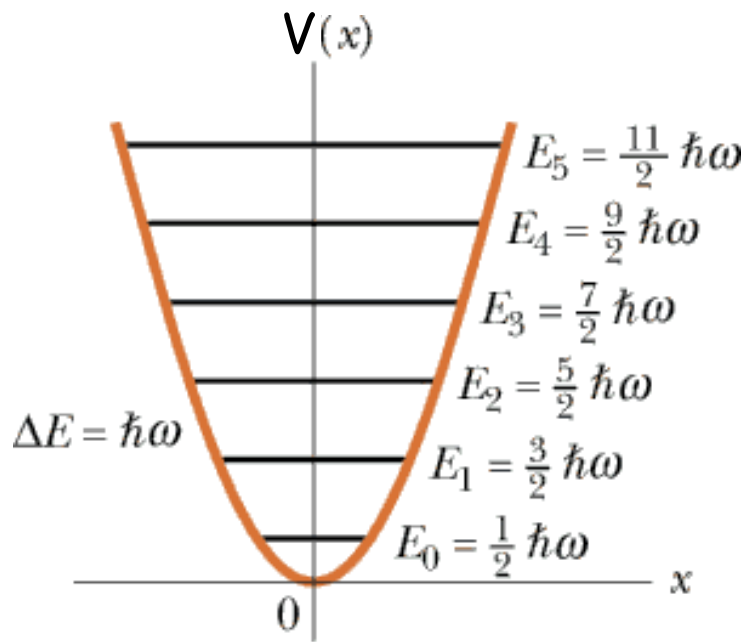
$$\underbrace{\hbar\omega(\hat{a}_+ \hat{a}_- + 1/2)}_{=0} \psi_0 = E_0 \psi_0$$

\hat{H} (Hamiltonian); BTW note how easy
was the calculation of E_0 .

Then:

$$E_0 = \frac{1}{2} \hbar\omega$$

>0 as expected. Deep
meaning: QM has "Zero point
energy". The harmonic
oscillators, or any other QM
problem, are never still!



Equally spaced levels

$$E_n = \left(n + \frac{1}{2} \right) \hbar \omega$$

Solutions are

$$\psi_n(x) = A_n (\hat{a}_+)^n \psi_0(x)$$

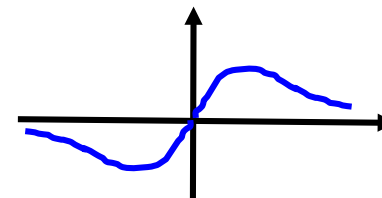


Example 2.4: construct state 1

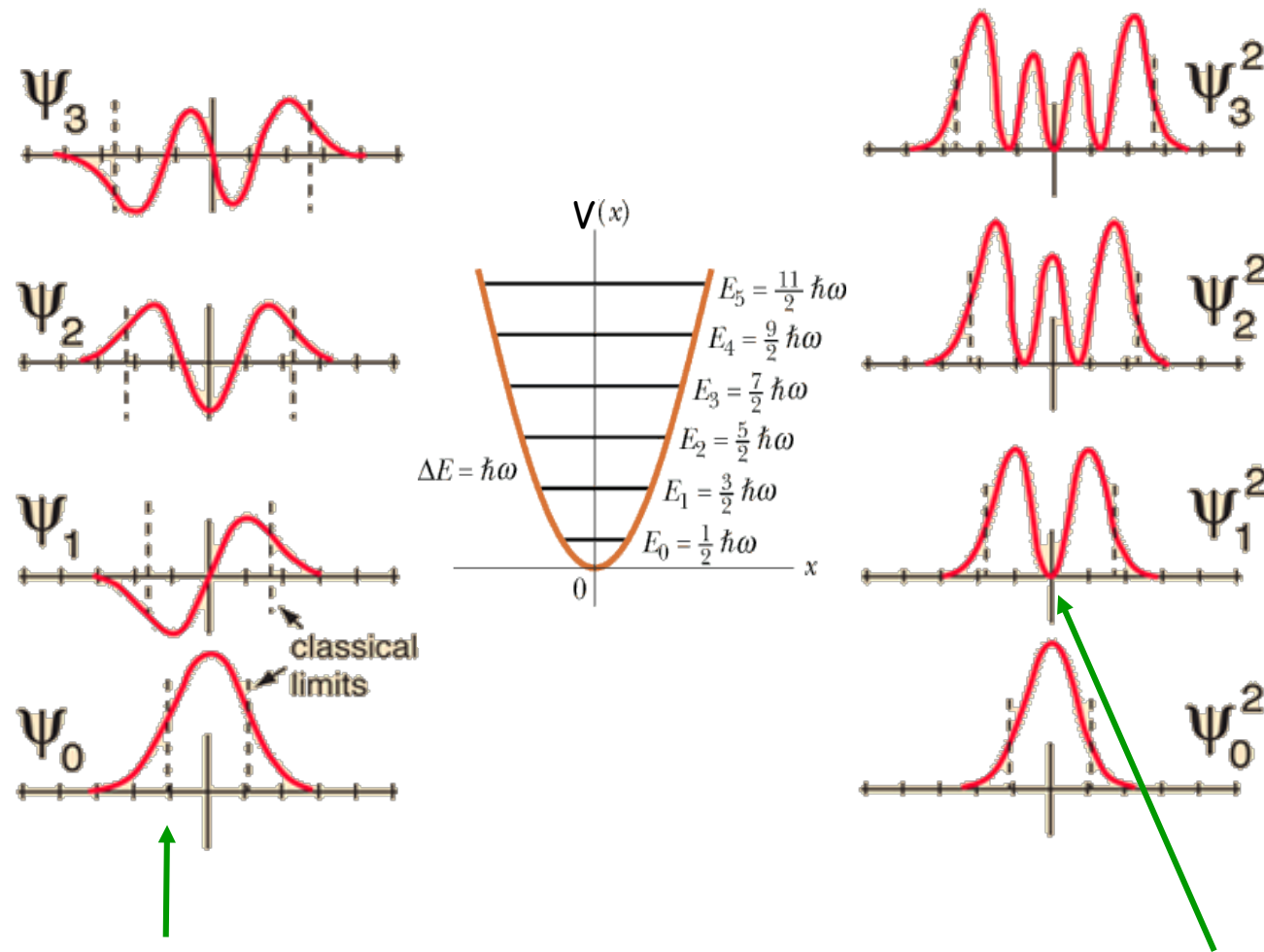
$$\psi_1(x) = A_1 \hat{a}_+ \psi_0 = \frac{A_1}{\sqrt{2\hbar m\omega}} \left(-\hbar \frac{d}{dx} + m\omega x \right) \left(\frac{m\omega}{\pi \hbar} \right)^{1/4} e^{-\frac{m\omega}{2\hbar} x^2}$$

$$= A_1 \left(\frac{m\omega}{\pi \hbar} \right)^{1/4} \sqrt{\frac{2m\omega}{\hbar}} x e^{-\frac{m\omega}{2\hbar} x^2}$$

Brings down an x



Odd function



Note: particle can be found outside the classical region

~Zero chance near nodes.

It can be shown, following the textbook, that:

$$\begin{aligned}\hat{a}_+ \psi_n &= \sqrt{n+1} \psi_{n+1} & \psi_n &= \frac{1}{\underbrace{\sqrt{n!}}_{A_n}} (\hat{a}_+)^n \psi_0 \\ \hat{a}_- \psi_n &= \sqrt{n} \psi_{n-1}\end{aligned}$$

$$\int_{-\infty}^{\infty} \psi_m^* \psi_n dx = \delta_{mn}$$

Orthonormal, like we found before for square well.


Example 2.5: find expectation value of V in the n -th state.

$$\langle V \rangle = \left\langle \frac{1}{2} m \omega^2 x^2 \right\rangle = \frac{1}{2} m \omega^2 \int_{-\infty}^{\infty} \psi_n^* x^2 \psi_n dx.$$

It may be tempting to write ψ_n as a Gaussian with some polynomial in front. However, there is a **simpler** way.

From $\hat{a}_{\pm} \equiv \frac{1}{\sqrt{2\hbar m \omega}} (\mp i \hat{p} + m \omega \hat{x})$ used before, deduce:

$$\hat{x} = \sqrt{\frac{\hbar}{2m\omega}} (\hat{a}_+ + \hat{a}_-); \quad \hat{p} = i \sqrt{\frac{\hbar m \omega}{2}} (\hat{a}_+ - \hat{a}_-).$$


$$\hat{x}^2 = \frac{\hbar}{2m\omega} \left[(\hat{a}_+)^2 + (\hat{a}_+ \hat{a}_-) + (\hat{a}_- \hat{a}_+) + (\hat{a}_-)^2 \right]$$

Creates ψ_{n-2} orthog to ψ_n

$$\langle V \rangle = \frac{\hbar\omega}{4} \int \psi_n^* \left[(\hat{a}_+)^2 + (\hat{a}_+\hat{a}_-) + (\hat{a}_-\hat{a}_+) + (\hat{a}_-)^2 \right] \psi_n dx.$$

Creates ψ_{n+2} orthog to ψ_n

$$\hat{a}_+\psi_n = \sqrt{n+1}\psi_{n+1}, \quad \hat{a}_-\psi_n = \sqrt{n}\psi_{n-1}$$

$$\hat{a}_-\hat{a}_+\psi_n = \hat{a}_-\sqrt{n+1}\psi_{n+1} = \sqrt{n+1}\underbrace{\hat{a}_-\psi_{n+1}}_{\sqrt{n+1}\psi_n} = (n+1)\psi_n$$

$$\langle V \rangle = \frac{\hbar\omega}{4} (n + n + 1) = \frac{1}{2}\hbar\omega \left(n + \frac{1}{2} \right) = \frac{1}{2} E_n$$

About HW3

(2.10) and (2.11)

$$\text{Using } \psi_m(x) = \frac{1}{\sqrt{m!}} (\hat{a}_+)^m \psi_0(x), \text{ and } \hat{a}_+ = \frac{1}{\sqrt{2\hbar m\omega}} (-i\hat{p} + m\omega\hat{x}),$$

$$\text{and } \psi_0(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} e^{-\frac{m\omega}{2\hbar}x^2},$$

you move the entire problem into "x space" and integrate, etc.

(2.12)

Following example 2.5 book, you keep wave functions in terms of \hat{a}_+ operators and, instead, transform, for example, \hat{x}^2 or \hat{p}^2 in the language of \hat{a} and \hat{a}_+ . Then, find expectation values. Often this procedure is easier.