

Consider now the **excited states**, starting with  $n=2$ :

$$E_2 = \frac{-13.6 \text{ eV}}{4}$$

$$= -3.4 \text{ eV}$$

Because  $n = j_{\max} + l + 1$ , then  $n=2$  allows for  $l=0, j_{\max}=1$  or  $l=1 (m=-1, 0, +1), j_{\max}=0$ . This will be the 2s and three 2p's orbitals. **Degeneracy 4** (in general, degeneracy is  $n^2$ ).

(1) If  $l=0, j_{\max}=1$ , use  $c_1 = [2(j + l + 1 - n) / (j+1)(j+2l+2)] c_0$  with  $n=2, j=0$  and  $l=0$ , and you get  $c_1 = -c_0$  so the polynomial becomes  $v(\rho) = c_0(1-\rho)$  with  $c_0$  again used to normalize.

$$R_{20}(r) = \frac{c_0}{2a} \left( 1 - \frac{r}{2a} \right) e^{-r/2a}$$

$$Y_0^0(\theta, \phi) = 1/\sqrt{4\pi}$$

The factors 2 arise from  $\rho = (1/na) r$ .

(2) If  $l=1, j_{max}=0$ , you get  $c_1 = 0$  so the polynomial is only  $c_0$  as it happens in the ground state.

Use the general formula  $R_{nl}(r) = \frac{1}{r} \rho^{l+1} e^{-\rho} v(\rho)$  with  $l=1$  and  $n=2$ .

$$R_{21}(r) = \frac{c_0}{4a^2} r e^{-r/2a}$$

The front factor comes from  $\rho^2/r = (r/2a)^2 / r$ .

The spherical harmonics are:

$$Y_1^0 = \left(\frac{3}{4\pi}\right)^{1/2} \cos \theta$$

$$Y_1^{\pm 1} = \mp \left(\frac{3}{8\pi}\right)^{1/2} \sin \theta e^{\pm i\phi}$$

In general, the polynomials  $v(\rho)$  are called **associated Laguerre polynomials**

$$v(\rho) = L_{n-l-1}^{2l+1}(2\rho)$$

and there are tables with these polynomials.

Putting all together, for arbitrary  $(n, l, m)$  the normalized wave functions are:

$$\psi_{nlm} = \sqrt{\left(\frac{2}{na}\right)^3 \frac{(n-l-1)!}{2n[(n+l)!]^3}} e^{-r/na} \left(\frac{2r}{na}\right)^l \left[ L_{n-l-1}^{2l+1}\left(\frac{2r}{na}\right) \right] Y_l^m(\theta, \phi)$$


$$R_{nl}(r)$$

$$R_{10} = 2a^{-3/2} \exp(-r/a)$$

$$R_{20} = \frac{1}{\sqrt{2}} a^{-3/2} \left(1 - \frac{1}{2} \frac{r}{a}\right) \exp(-r/2a)$$

$$R_{21} = \frac{1}{\sqrt{24}} a^{-3/2} \frac{r}{a} \exp(-r/2a)$$

$$R_{30} = \frac{2}{\sqrt{27}} a^{-3/2} \left(1 - \frac{2}{3} \frac{r}{a} + \frac{2}{27} \left(\frac{r}{a}\right)^2\right) \exp(-r/3a)$$

$$R_{31} = \frac{8}{27\sqrt{6}} a^{-3/2} \left(1 - \frac{1}{6} \frac{r}{a}\right) \left(\frac{r}{a}\right) \exp(-r/3a)$$

$$R_{32} = \frac{4}{81\sqrt{30}} a^{-3/2} \left(\frac{r}{a}\right)^2 \exp(-r/3a)$$

$$R_{40} = \frac{1}{4} a^{-3/2} \left(1 - \frac{3}{4} \frac{r}{a} + \frac{1}{8} \left(\frac{r}{a}\right)^2 - \frac{1}{192} \left(\frac{r}{a}\right)^3\right) \exp(-r/4a)$$

$$R_{41} = \frac{\sqrt{5}}{16\sqrt{3}} a^{-3/2} \left(1 - \frac{1}{4} \frac{r}{a} + \frac{1}{80} \left(\frac{r}{a}\right)^2\right) \frac{r}{a} \exp(-r/4a)$$

$$R_{42} = \frac{1}{64\sqrt{5}} a^{-3/2} \left(1 - \frac{1}{12} \frac{r}{a}\right) \left(\frac{r}{a}\right)^2 \exp(-r/4a)$$

$$R_{43} = \frac{1}{768\sqrt{35}} a^{-3/2} \left(\frac{r}{a}\right)^3 \exp(-r/4a)$$

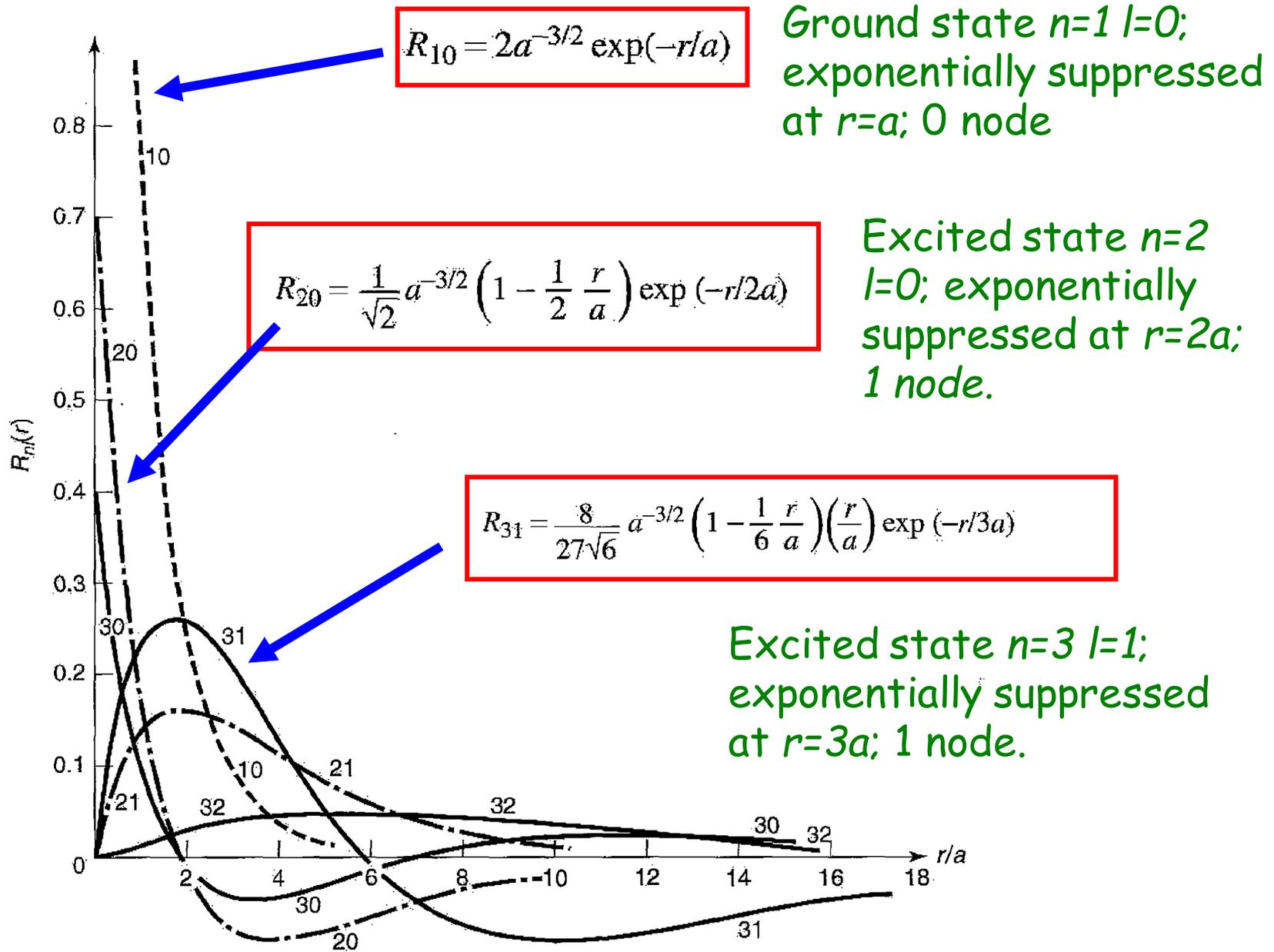
Only  $l=0$  are nonzero at the center  $r=0$ , like in the spherical well.

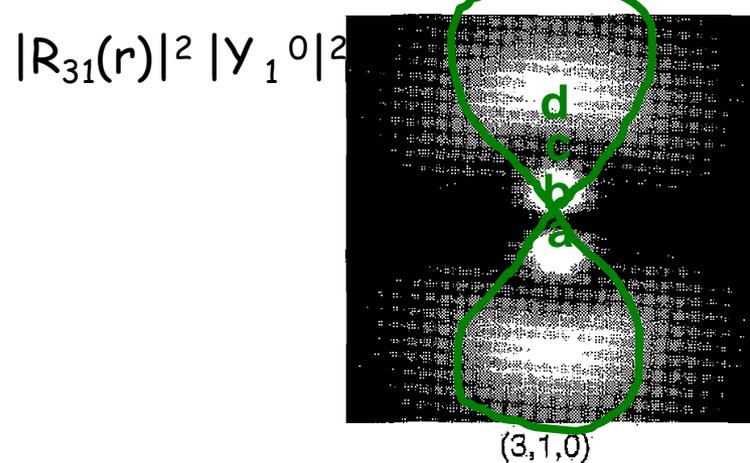
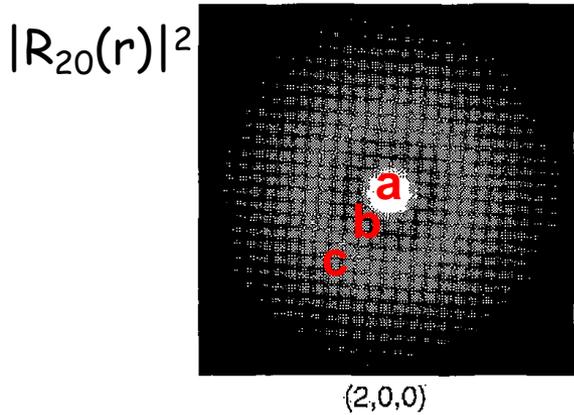
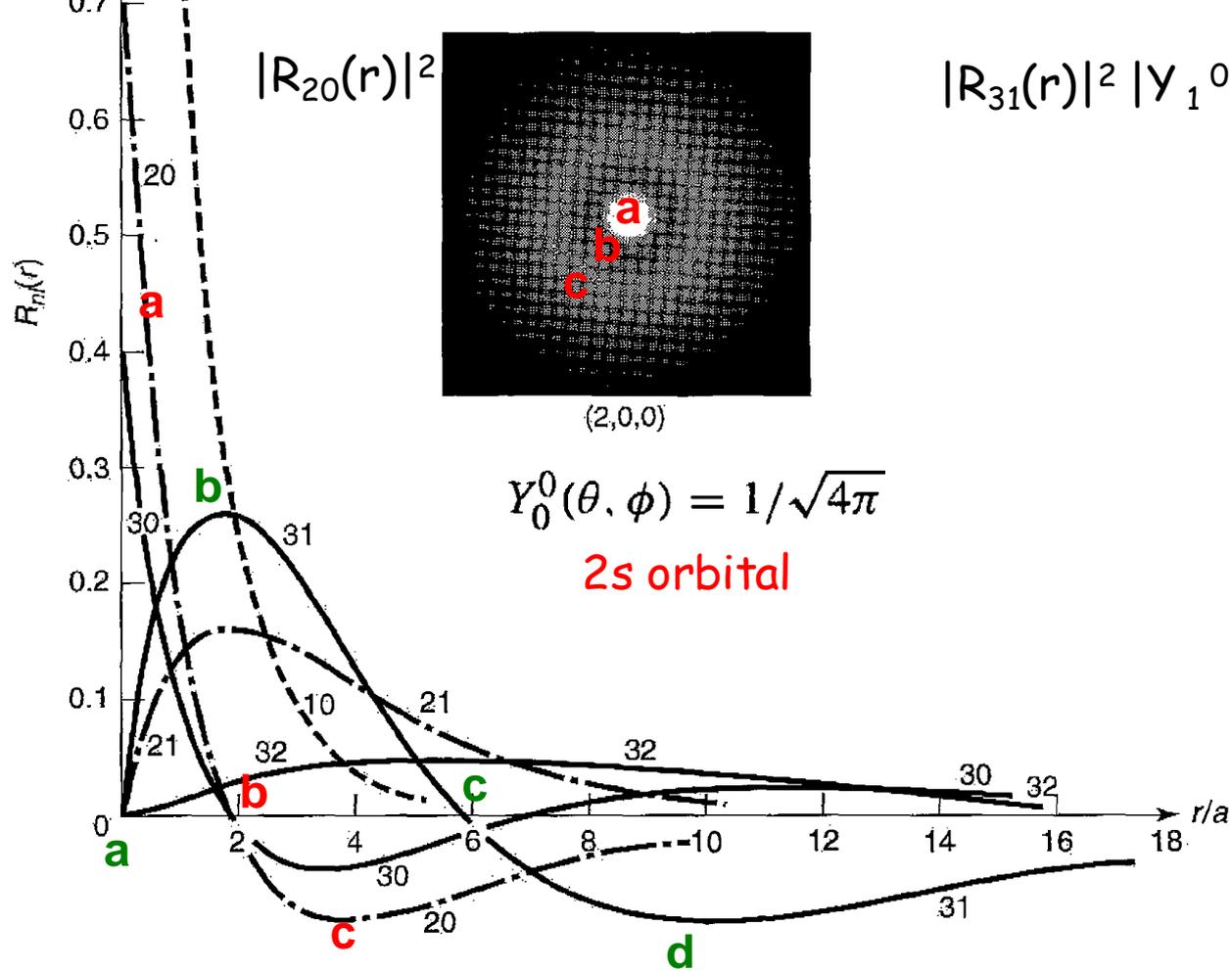
3 nodes

2 nodes

1 nodes

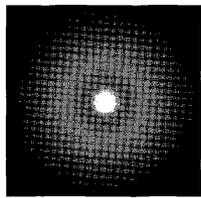
0 nodes





$$Y_1^0 = \left(\frac{3}{4\pi}\right)^{1/2} \cos \theta$$

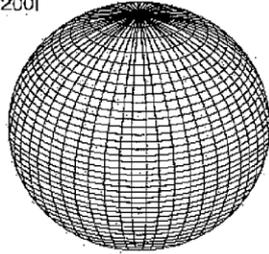
3p<sub>z</sub> orbital



(2,0,0)

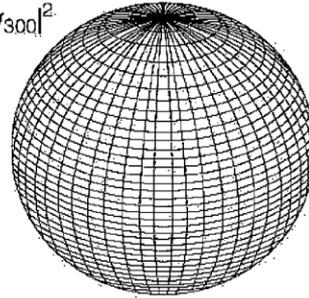
2s orbital

$|\psi_{200}|^2$

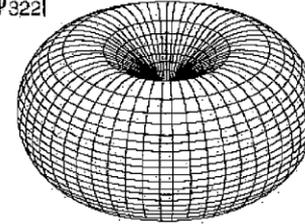


3s orbital

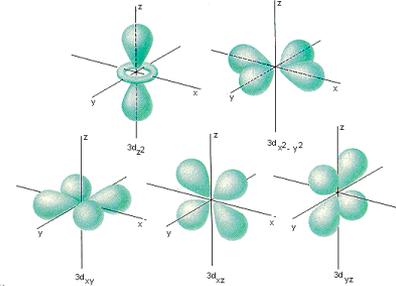
$|\psi_{300}|^2$



$|\psi_{322}|^2$

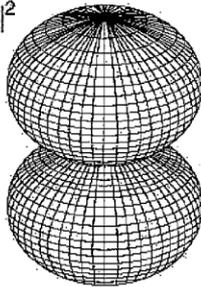


linear combination  
of canonical 3d  
orbitals:  $x^2-y^2$ ,  
 $3z^2-r^2$ ,  $xy, yz, zx$



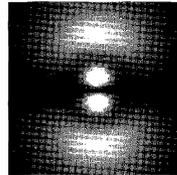
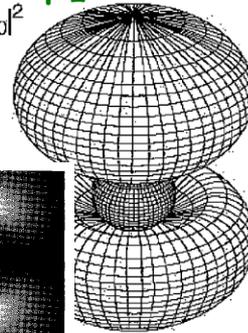
2p<sub>z</sub> orbital

$|\psi_{210}|^2$



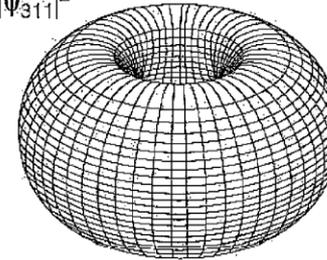
3p<sub>z</sub> orbital

$|\psi_{310}|^2$



(3,1,0)

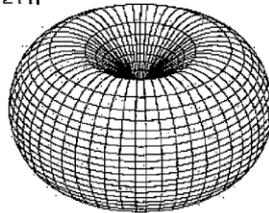
$|\psi_{311}|^2$



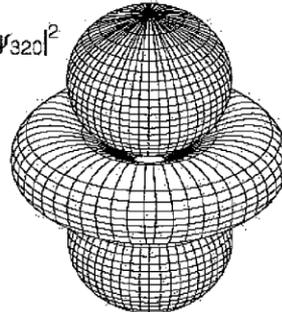
3(p<sub>x</sub>+ip<sub>y</sub>)  
orbital

2(p<sub>x</sub>+ip<sub>y</sub>)  
orbital

$|\psi_{211}|^2$

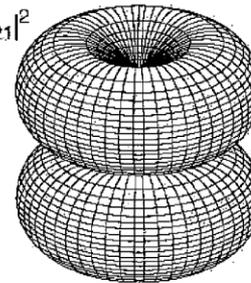


$|\psi_{320}|^2$



3d<sub>3z^2-r^2</sub> orbital

$|\psi_{321}|^2$



linear combination  
of canonical 3d  
orbitals

And as usual, the wave functions are **orthonormal**:

$$\int \psi_{nlm}^* \psi_{n'l'm'} r^2 \sin \theta dr d\theta d\phi = \delta_{nn'} \delta_{ll'} \delta_{mm'}$$

Because of  
radial  
equation.

Because of  
spherical  
harmonics.

## 4.2.2: The Spectrum of Hydrogen

Emission of photons:

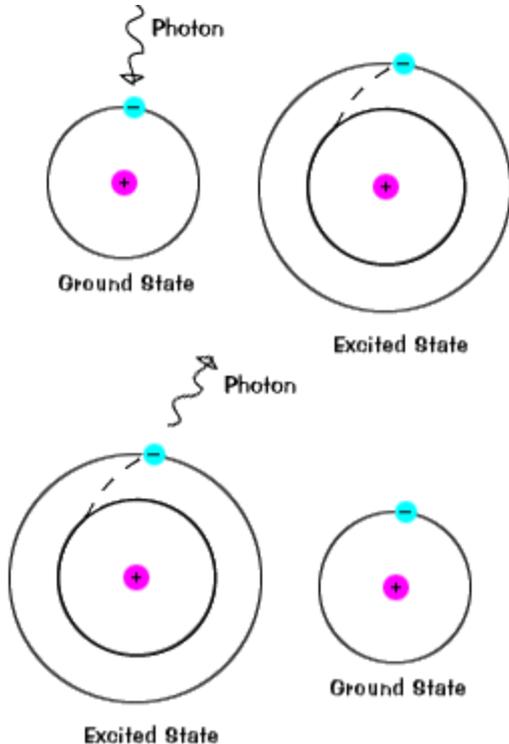
$$E_\gamma = E_i - E_f = -13.6 \text{ eV} \left( \frac{1}{n_i^2} - \frac{1}{n_f^2} \right)$$

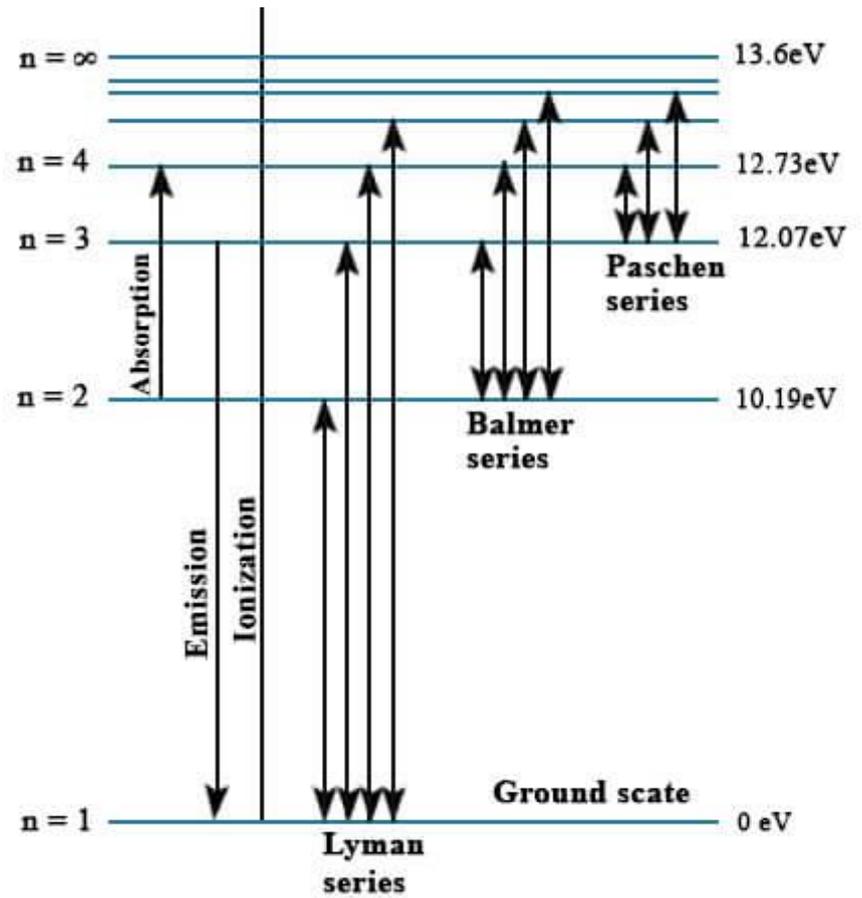
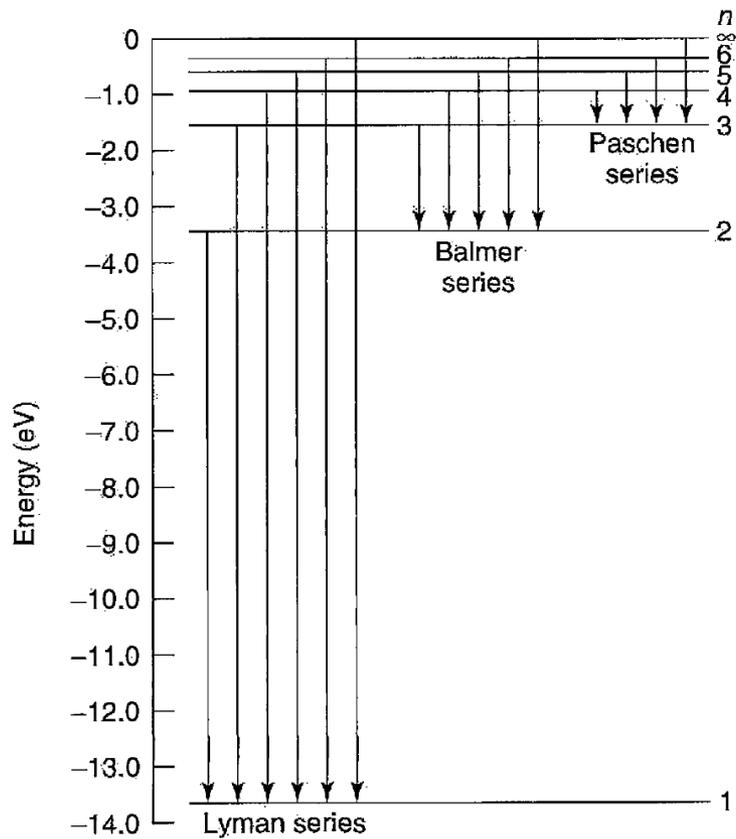
$$E_\gamma = h\nu \quad \lambda = c/\nu$$

$$\frac{1}{\lambda} = R \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

$$R = \frac{m}{4\pi c \hbar^3} \left( \frac{e^2}{4\pi \epsilon_0} \right)^2 = 1.097 \times 10^7 \text{ m}^{-1}$$

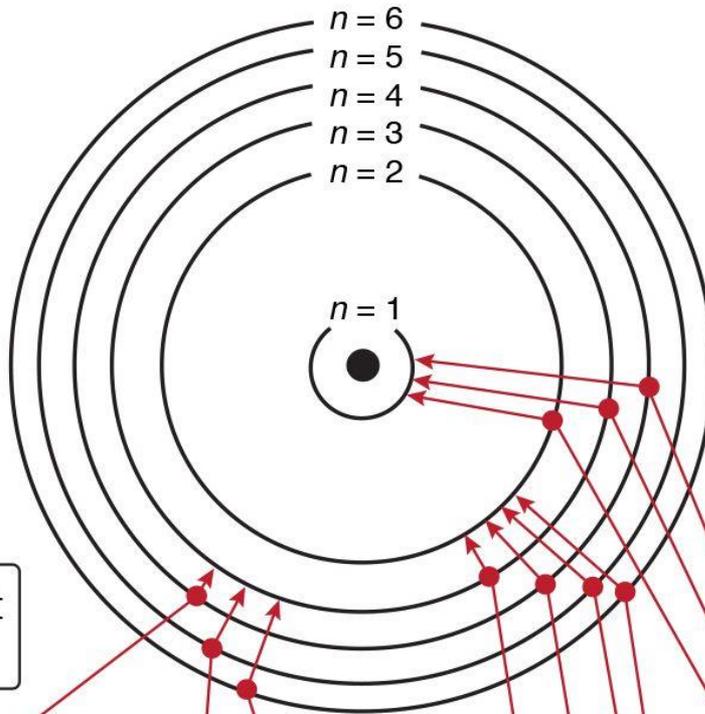
R= Rydberg constant; formula found before Sch Eq just empirically.



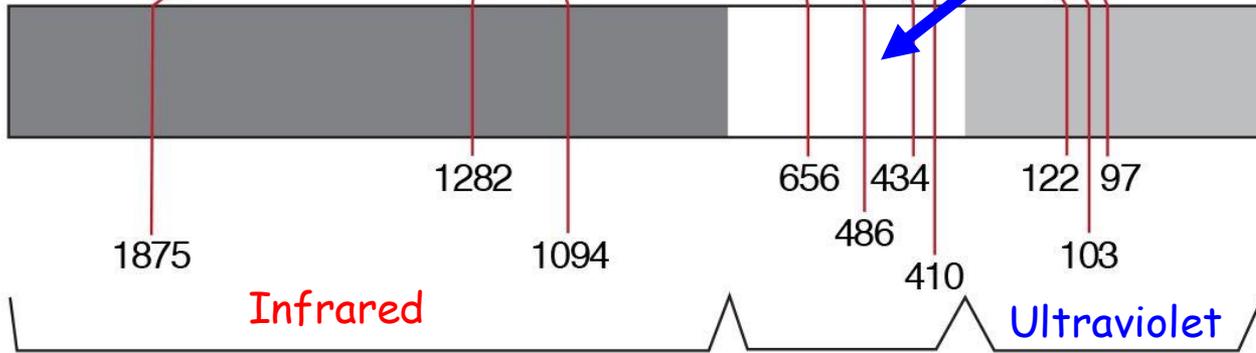
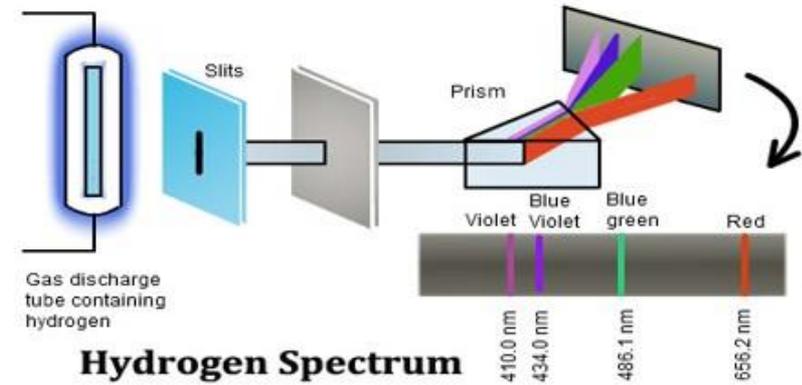


# Bohr Model for Hydrogen Atom

(measurement in nanometers)



UV = Ultraviolet  
IR = Infrared



X rays  $\gamma$  rays ...

<- Microwave IR; Paschen,  $n=3$  Visible; Balmer,  $n=2$  UV; Lyman,  $n=1$