<u>6.5 Zeeman Effect:</u>

Electron in an atom in the presence of a magnetic field. Before only "spin" S in a magnetic field was considered. Now we add the orbital angular momentum L because electron is orbiting the proton.

$$H'_{Z} = -(\mu_{l} + \mu_{s}) \cdot \mathbf{B}_{ext}$$

$$\mu_{s} = -\frac{e}{m}\mathbf{S} \qquad \mu_{l} = -\frac{e}{2m}\mathbf{L} \qquad \text{Missing 2 in spin is due to relativity.}$$

$$H'_{Z} = \frac{e}{2m}(\mathbf{L} + 2\mathbf{S}) \cdot \mathbf{B}_{ext} \qquad \mathbf{B}_{ext} \quad \mathbf{B}_{in} = \frac{1}{4\pi\epsilon_{0}}\frac{e}{mc^{2}r^{3}}\mathbf{L}$$

$$B_{\rm ext} \ll B_{\rm int}$$

If magnetic field is very small, then the fine structure constant must be considered as part of the H_0 before adding magnetic field.

$$B_{\rm ext} \gg B_{\rm int}$$

If magnetic field not too small, then the fine structure constant is neglected in H_0 before adding magnetic field. (1) $B_{\text{ext}} \ll B_{\text{int}}$ Fine structure is important. Good quantum numbers are (n, l, s, j, m_i)

$$E_Z^1 = \langle nljm_j | H'_Z | nljm_j \rangle = \frac{e}{2m} \mathbf{B}_{\text{ext}} \cdot \langle \mathbf{L} + 2\mathbf{S} \rangle$$

It can be shown that:

$$\langle \mathbf{L} + 2\mathbf{S} \rangle = \left[1 + \frac{j(j+1) - l(l+1) + 3/4}{2j(j+1)} \right] \langle \mathbf{J} \rangle$$

Lande' g_J factor

Consider magnetic field along z axis:

Then:
$$E_Z^1 = \mu_B g_J B_{\text{ext}} m_j$$

Simple split linear with B_{ext} : some levels up, others down.

where the Bohr magneton is defined as:

$$\mu_B \equiv \frac{e\hbar}{2m} = 5.788 \times 10^{-5} \text{ eV/T}$$

Example: ground state has n=1, l=0, s=1/2, j=1/2, m_j = $\pm 1/2$ (g_J=2)

Egs =
$$-13.6 \text{ eV}(1 + \alpha^2/4) \pm \mu_B B_{\text{ext}}$$



(2) $B_{\rm ext} \gg B_{\rm int}$

Magnetic field dominates. Good quantum numbers are (n, l, s, m_l, m_s) because magnetic field is larger than fine structure correction.

$$E_{nm_lm_s} = -\frac{13.6 \text{ eV}}{n^2} + \mu_B B_{\text{ext}}(m_l + 2m_s)$$
$$E_{\text{fs}}^1 = \langle nl \ m_l \ m_s | (H'_r + H'_{\text{so}}) | nl \ m_l \ m_s \rangle$$
$$E_{\text{fs}}^1 = \frac{13.6 \text{ eV}}{n^3} \alpha^2 \left\{ \frac{3}{4n} - \left[\frac{l(l+1) - m_lm_s}{l(l+1/2)(l+1)} \right] \right\}$$



Chapter 7: The variational principle

This is a common occurrence. Suppose you have a Hamiltonian that (i) cannot be solved exactly and (ii) where perturbation theory cannot be applied because there is no simple H_0 and/or because there is no small H'.

Then, what do we do? \otimes

One possibility is to use the variational principle: it does not give you the exact answer but gives you an upper bound that is often sufficient.

Select any wave function you wish. Call it Ψ . The claim is that always:

$$E_{\rm gs} \leq \langle \psi | H | \psi \rangle \equiv \langle H \rangle$$

Although we do not know explicitly the eigenstates of H, because we cannot solve the problem exactly, we know they exist.

$$H\psi_n=E_n\psi_n$$

Then, we can expand our proposed variational wave function in the complete basis of eigenstates:

$$\psi = \sum_n c_n \psi_n$$

If Ψ is normalized, then:

$$1 = \langle \psi | \psi \rangle = \left\langle \sum_{m} c_{m} \psi_{m} \right| \sum_{n} c_{n} \psi_{n} \right\rangle = \sum_{m} \sum_{n} c_{m}^{*} c_{n} \langle \psi_{m} | \psi_{n} \rangle = \sum_{n} |c_{n}|^{2}$$

Repeating with *H* included, we find:

$$\langle H \rangle = \left\langle \sum_{m} c_{m} \psi_{m} \middle| H \sum_{n} c_{n} \psi_{n} \right\rangle = \sum_{m} \sum_{n} c_{m}^{*} E_{n} c_{n} \langle \psi_{m} | \psi_{n} \rangle = \sum_{n} E_{n} |c_{n}|^{2}$$

But the ground state has the lowest energy by definition: $E_{gs} \leq E_n$. Then: $\langle H \rangle \geq E_{gs} \sum |c_n|^2 = E_{gs}$

The variational principle is powerful, easy to use, and accurate if you have a good intuition on how the wave function should look like. Problem: you do NOT know how close your result is compared to the exact result. You only know you are above.

Example 7.1:

Consider the 1D Harmonic Oscillator with H:

$$H = -\frac{\hbar^2}{2m}\frac{d^2}{dx^2} + \frac{1}{2}m\omega^2 x^2$$

Here we know the answer exactly, but we pretend we do not.

As a "trial" wave function we will use a Gaussian exponential. Using Gaussians is very common, because the integrals are known.

$$\psi(x) = Ae^{-bx^2}$$

A is the normalization and b is called a "variational parameter" that we will optimize by minimizing the energy.

Normalization:

$$|\psi(x)|^{2}$$

$$1 = |A|^{2} \int_{-\infty}^{\infty} e^{-2bx^{2}} dx = |A|^{2} \sqrt{\frac{\pi}{2b}} \Rightarrow A = \left(\frac{2b}{\pi}\right)^{1/4}$$

Next we need the expectation value of the Hamiltonian:

$$\langle H \rangle = \langle T \rangle + \langle V \rangle$$

For the kinetic energy:

$$\langle T \rangle = -\frac{\hbar^2}{2m} |A|^2 \int_{-\infty}^{\infty} e^{-bx^2} \frac{d^2}{dx^2} \left(e^{-bx^2} \right) dx = \frac{\hbar^2 b}{2m}$$

For the potential energy:

$$\langle V \rangle = \frac{1}{2}m\omega^2 |A|^2 \int_{-\infty}^{\infty} e^{-2bx^2} x^2 dx = \frac{m\omega^2}{8b}$$

Adding kinetic and potential energy: $\langle H \rangle = \frac{\hbar^2 b}{2m} + \frac{m\omega^2}{8b}$

Let us now "optimize" the "variational parameter"

$$\frac{d}{db}\langle H\rangle = \frac{\hbar^2}{2m} - \frac{m\omega^2}{8b^2} = 0 \implies b = \frac{m\omega}{2\hbar}$$

If we introduce the "optimal b" into <H>, we obtain:

$$\langle H \rangle_{\min} = \frac{1}{2} \hbar \omega$$

which is the exact result, by chance, in this simple example. In the vast majority of cases, you will **not** find the exact result.