

10/29

Last time:

Properties of solids.

Interacting atoms.



$$x_{ij} = x_i - x_j \quad 3N \text{ coordinates}$$

$$H = V_0 + \frac{1}{2} \sum_{i,j} m_i \dot{x}_{ij}^2 + \frac{1}{2} \sum_{i,j} A_{ij} x_{ij}^2$$

↓ change of basis to q_r (generalized coordinates),

$$H = V_0 + \frac{1}{2} \sum_{r=1}^{3N} (\dot{q}_r^2 + \omega_r^2 q_r^2)$$

We see that H is written in terms of $3N$ Hamiltonians \tilde{H}_r

$$\tilde{H}_r = \frac{1}{2} (\dot{q}_r^2 + \omega_r^2 q_r^2) \quad \text{1D harmonic oscillator.}$$

We know that

$$E_r = (n_r + \frac{1}{2}) \hbar \omega_r \quad n_r = 0, 1, \dots$$

The state of the system is given by $3N$ integers $n_r : \{n_1, n_2, \dots, n_{3N}\}$

We can associate the occupation numbers n_r for each oscillator with a number of "quasiparticles" like we did with photons that are the quasiparticles associated to electromagnetic waves. The quasiparticles associated to lattice vibrations or sound waves are called **phonons**.

Then

$$\begin{aligned}
 E_{n_1, n_2, \dots, n_{3N}} &= V_0 + \sum_{r=1}^{3N} (n_r + \frac{1}{2}) \hbar \omega_r = \\
 &= -N\gamma + \sum_{r=1}^{3N} n_r \hbar \omega_r
 \end{aligned}$$

Where

$$-N\eta = V_0 + \frac{1}{2} \sum_r \hbar \omega_r$$

energy of the
system with
all $n_r = 0$.

η represents the binding energy per atom when they are at rest at $T=0$. Notice that it is not just V_0 but it contains also the $T=0$ energy of the oscillators.

Partition function:

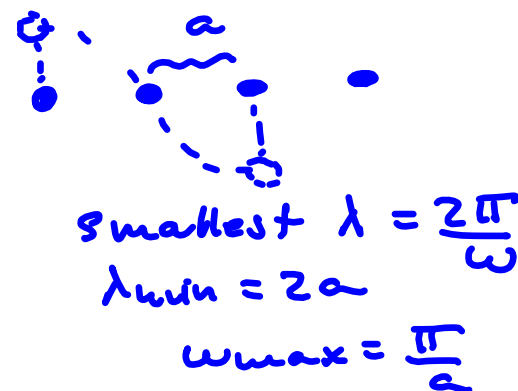
Phonons are bosons since there are no restrictions on the value of n_r .

$$\begin{aligned}
 Z &= \sum_{n_1, n_2, \dots, n_{3N}} e^{-\beta(-N\gamma + n_1 \hbar \omega_1 + \dots + n_{3N} \hbar \omega_{3N})} \\
 &= e^{+\beta N \gamma} \underbrace{\left(\sum_{n_1=0}^{\infty} e^{-\beta n_1 \hbar \omega_1} \right)}_{\frac{1}{1 - e^{-\beta \hbar \omega_1}}} \dots \underbrace{\left(\sum_{n_{3N}=0}^{\infty} e^{-\beta n_{3N} \hbar \omega_{3N}} \right)}_{\frac{1}{1 - e^{-\beta \hbar \omega_{3N}}}}
 \end{aligned}$$

$$z = e^{\beta N \gamma} \prod_{r=1}^{3N} \frac{1}{1 - e^{-\beta \hbar \omega_r}}$$

$$\ln z = \beta N \gamma - \sum_{r=1}^{3N} \ln(1 - e^{-\beta \hbar \omega_r})$$

$$\omega_{\max} = \frac{F}{a} = \frac{F}{R_0}$$



Experimental slope of density of normal frequencies

If we assume that there are many ω 's in an interval $d\omega$ then we can write

$$\ln Z = \beta N \gamma - \int_0^{\infty} \ln(1 - e^{-\beta \hbar \omega}) \sigma(\omega) d\omega$$

$$\therefore \bar{E} = -\frac{\partial \ln Z}{\partial \beta} = -N \gamma + \int_0^{\infty} \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1} \sigma(\omega) d\omega$$

$$C_V = \left. \frac{\partial \bar{E}}{\partial T} \right|_V = -k \beta^2 \left. \frac{\partial \bar{E}}{\partial \beta} \right|_V = k \int_0^{\infty} \frac{e^{\beta \hbar \omega} (\beta \hbar \omega)^2 \sigma(\omega) d\omega}{(e^{\beta \hbar \omega} - 1)^2}$$

Hardest part of the problem is to find $\sigma(\omega)$.

First consider some limits:

$\frac{\hbar\omega}{kT}$ defines the temperature regime.

If $\frac{\hbar\omega}{kT} \ll 1 \Rightarrow \hbar\omega \ll kT$ and T is large.

Then $\frac{\hbar\omega_{\max}}{kT} \ll 1$ Notice that $\sigma(\omega) = 0$ for $\omega > \omega_{\max}$

and $\frac{\hbar\omega_r}{kT} \ll 1 \quad \forall r$

$$\therefore e^{\beta \hbar\omega} \approx 1 + \beta \hbar\omega + \dots$$

Then

$$C_V = k \int_0^\infty \frac{e^{\beta \hbar \omega} (\beta \hbar \omega)^2 \sigma(\omega) d\omega}{(e^{\beta \hbar \omega} - 1)^2} =$$

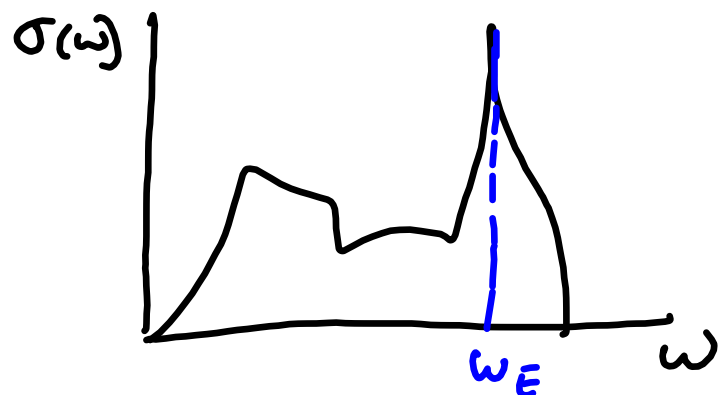
$$\cong k \int_0^\infty \sigma(\omega) d\omega = k 3N$$

Dulong and
Petit law
for C_V of a
solid at
room temperature

For $\frac{\hbar \omega}{kT} \gg 1$ (low T) we

need to make an assumption about $\sigma(\omega)$.

1) we saw time ago Einstein's assumption:



$$g_E(\omega) = \delta(\omega - \omega_E)$$

$$C_V = k \int_0^{\infty} \frac{e^{-\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2} (\beta \hbar \omega)^2 \delta(\omega - \omega_E) d\omega =$$

$$= \frac{k e^{-\beta \hbar \omega_E}}{(e^{\beta \hbar \omega_E} - 1)^2} (\beta \hbar \omega_E)^2 \longrightarrow k e^{-\beta \hbar \omega_E} (\beta \hbar \omega_E)^2$$

$\frac{\hbar \omega_E}{kT} \gg 1$

goes to zero
like an
exponential.

But experimentally it is known that

$$C_V \propto T^3$$

To improve the agreement we will use

2) Debye's approximation.

He viewed the solid as a continuous elastic medium. But with the assumption that there is a minimum λ that can propagate due to the finite separation between atoms.

$\sigma_c(\omega)$ der is calculated in the same way
 as for photons - we found the number of
 modes per unit frequency for a particle with
 energy $\hbar\omega$ in a box with periodic
 boundary conditions. For photons we multiplied
 that number by 2 to account for the 2
 polarizations. Now for phonons we need to
 multiply by 3 since there are 3 possible
 propagation directions.

$$\sigma_c(\omega) d\omega = 3 \frac{V}{2\pi^2 c_s^3} \omega^2 d\omega$$

If $c_l \neq c_t$ then $\frac{1}{c_s^3} = \frac{1}{c_l^3} + \frac{2}{c_t^3}$

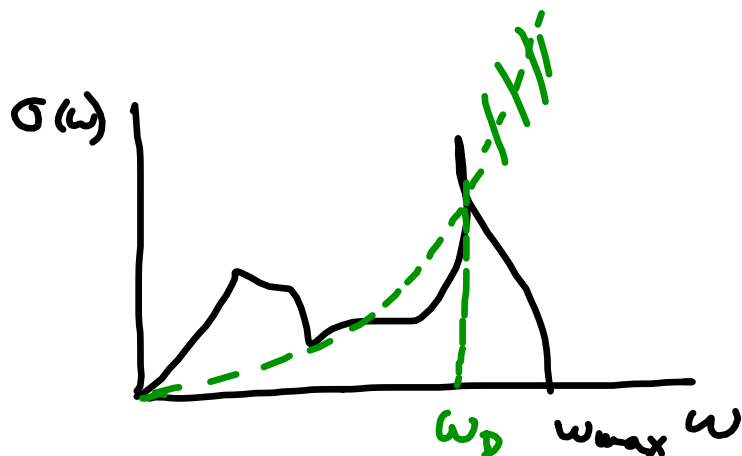
c_l : longitudinal speed of sound

c_t : transversal " " " " .

Debye's approximation:

$$\sigma_D(\omega) = \begin{cases} \sigma_c(\omega) & \text{for } \omega \leq \omega_D \\ 0 & \text{for } \omega > \omega_D \end{cases}$$

ω_D : Debye's
Frequency



ω_D is obtained from the normalization condition:

$$3N = \int_0^{\infty} \sigma_D(\omega) d\omega = \int_0^{\omega_D} \sigma_c(\omega) d\omega = \frac{3V}{2\pi^2 c_s^3} \int_0^{\omega_D} \omega^2 d\omega =$$

$$= \frac{V}{2\pi^2 c_s^3} \omega_D^3 \Rightarrow \boxed{\omega_D = c_s \left(6\pi^2 \frac{N}{V} \right)^{1/3}}$$

Now we can calculate C_V for $\hbar\omega \gg kT$

$$C_V = k \int_0^{\omega_D} \frac{e^{-\beta \hbar \omega} (\beta \hbar \omega)^2 3V \omega^2 d\omega}{2\pi^2 c_s^3 (e^{-\beta \hbar \omega} - 1)^2}$$

Define $x = \beta \hbar \omega$

$$C_V = \frac{3Vk}{2\pi^2 (c_s \beta \hbar)^3} \int_0^{\beta \hbar \omega_D} \frac{e^{-x}}{(e^x - 1)^2} x^4 dx =$$

$$= 3Nk \frac{3k^3 T^3}{\hbar^3 \omega_D^3} \int_0^{\frac{\hbar \omega_D}{kT}} \frac{e^{-x}}{(e^x - 1)^2} x^4 dx =$$

$y = \frac{\hbar \omega_D}{kT}$

$$C_V = 3Nk \frac{3}{y^3} \int_0^y \frac{e^x}{(e^x - 1)^2} x^2 dx =$$

$$= 3Nk f_D\left(\frac{\hbar \omega_D}{kT}\right) = 3Nk f_D\left(\frac{\Theta_D}{T}\right)$$

Debye's function:

$$f_D(y) = \frac{3}{y^3} \int_0^y \frac{e^x}{(e^x - 1)^2} x^2 dx$$

$$y = \frac{\hbar \omega_D}{kT} = \frac{\Theta_D}{T}$$

with

$$\Theta_D = \frac{\hbar \omega_D}{k}$$

Debye's temperature.

material dependent.

Limits of $J_0(y)$:

If $y \ll 1 \Rightarrow (0 \ll T)$ in just T

$$e^x \approx 1+x$$

$$\frac{3}{y^3} \int_0^y \frac{e^x}{(e^x-1)^2} x^4 dx = \frac{3}{y^3} \int_0^y \frac{1}{(1+x-1)^2} x^4 dx =$$

$$= \frac{3}{y^3} \int_0^y x^2 dx = \frac{1}{y^3} x^3 \Big|_0^y = 1$$

Then for $\theta_D \ll T$

$$C_V = 3NK \quad (\text{Dulong and Petit})$$

For $y \gg 1$ ($\theta_D \gg T$) \Rightarrow low T

$$\frac{3}{\gamma^3} \int_0^{\infty} \frac{e^{-x} x^4}{(e^x - 1)^2} dx =$$

$\underbrace{\hspace{10em}}_{\frac{4\pi^4}{5}}$

We can integrate
up to ∞ because

$e^{-x} \ll 1$ for large w
due to large ρ .

$$\frac{4\pi^4}{5} \frac{3}{\theta_D^3} T^3 \propto T^3$$

In agreement with the
experiments.

Non-ideal classical gas:

Note: we will follow Reif but after that we will use the cluster expansion method in Kardar's book.

Assume N identical particles, m , V , T

T is high.

$$\frac{N}{V} = n \sim \text{small}$$

However, T is not as high or/and n is not so small that the ideal gas eq. of state can be used.

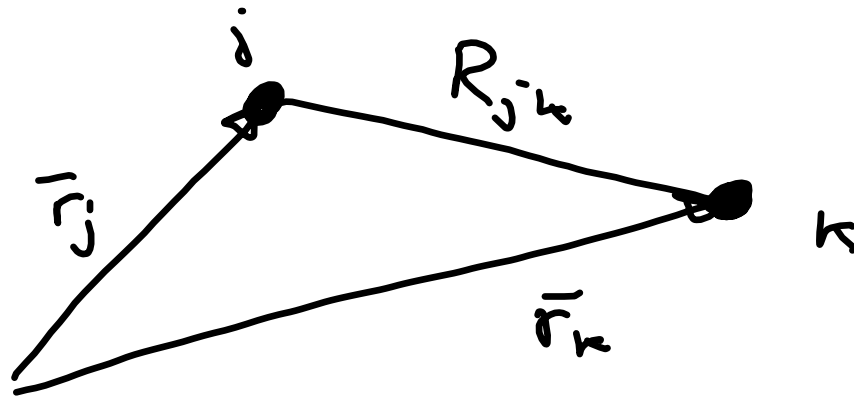
Now we will have to consider the interaction U between molecules that we neglected when we studied the ideal gas.

$$\tilde{H}_6 = K + U$$

$$K = \frac{1}{2m} \sum_{j=1}^N \vec{p}_j^2$$

classical limit of
the quantum case.

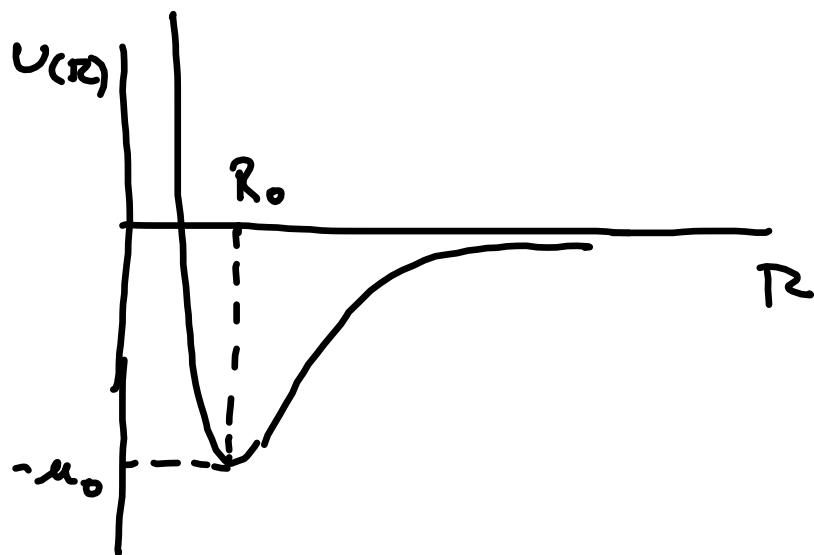
We will consider that only the interactions between 2 particles are relevant.



$$R_{jk} = |\vec{r}_j - \vec{r}_k|$$

$$u_{jk} = u(R_{jk})$$

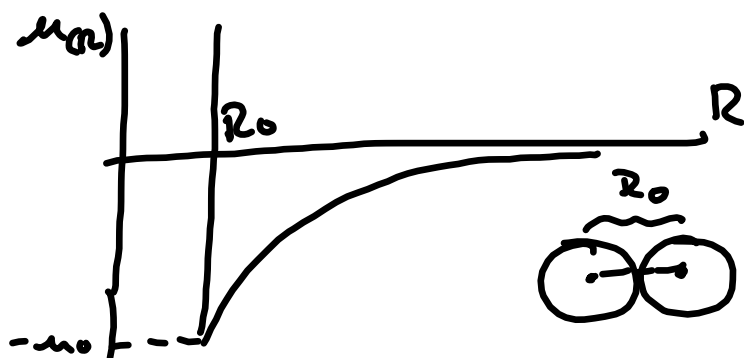
$$\begin{aligned}
 U &= u_{12} + u_{13} + \dots + u_{1N} + u_{23} + u_{24} + \dots = \\
 &= \sum_{\substack{j=1 \\ j < k}}^N \sum_{k=2}^N u_{jk} = \frac{1}{2} \sum_{j=1}^N \sum_{\substack{k=1 \\ j \neq k}}^N u_{jk}
 \end{aligned}$$



Lennard-Jones potential

$$u(R) = u_0 \left[\left(\frac{R_0}{R} \right)^{12} - 2 \left(\frac{R_0}{R} \right)^6 \right]$$

It can be simplified (van der Waals):



$$u(R) = \begin{cases} \infty & \text{for } R < R_0 \\ -u_0 \left(\frac{R_0}{R} \right)^s & R > R_0 \end{cases}$$

molecules with diameter radius $R_0/2$.
 $s > 3$ $s \approx 6$