

8/25

Last time:

Ideal Classical gas (high T
low n)

Quantum approach for energy levels:

$$S_q = Nk \ln V + \frac{3Nk}{2} \ln \bar{E} + \frac{3}{2} Nk \left[1 + \ln \frac{4\pi m}{3Nh} \right]$$

and with classical approach:

$$S_{class} = Nk \ln V + \frac{3Nk}{2} \ln \bar{E} + \text{constant}$$

In both approaches:

*calculated classically
assuming distinguishable particles*

$$\Omega \propto \frac{(2m\bar{E})^{\frac{3N}{2}} V^N}{(2m\bar{E} V^{2/3})^{\frac{3N}{2}}}$$

From S we can obtain:

$$\frac{1}{T} = \left. \frac{\partial S}{\partial E} \right|_{N, V} = \frac{3Nk}{2E} \Rightarrow E = \frac{3}{2} NkT$$

$$\therefore C_V = \left. \frac{\partial E}{\partial T} \right|_{V, N} = \frac{3}{2} Nk$$

$$\ln E = (S - kN \ln V - C) \frac{2}{3Nk}$$

$$E = C' e^{\frac{2S}{3Nk}} V^{-2/3}$$

$$P = - \left. \frac{\partial E}{\partial V} \right|_{N, S} = \frac{2}{3} \frac{E}{V} = \frac{2}{3} \frac{3}{2} \frac{kTN}{V} \Rightarrow PV = NkT$$

$$C_P = \left. \frac{\partial H}{\partial T} \right|_{P, N} = \left. \frac{\partial (E + PV)}{\partial T} \right|_{P, N} = \frac{\partial}{\partial T} \left(\frac{3}{2} NkT + NkT \right) = \frac{5}{2} Nk$$

$$\frac{C_p}{C_v} = \frac{5}{3} = \gamma$$

Isothermic process (reversible)

if $T = \text{constant} \Rightarrow E = \text{constant}$

$N = \text{constant}$

Then $P = \frac{n k T}{V} \propto \frac{1}{V}$ Boyle's law.

then $\Delta S = S_f - S_i = N k \ln V_f / V_i$

Gibbs paradox:

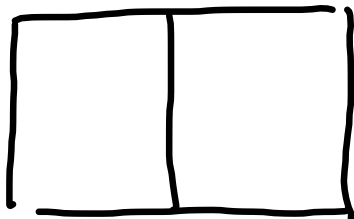
We found that

$$S(N, V, \bar{E}) = Nk \ln \left[\frac{V}{h^3} \left(\frac{4\pi m \bar{E}}{3N} \right)^{3/2} \right] + \frac{3Nk}{2} \quad (1)$$

If $N \rightarrow \alpha N$, $V \rightarrow \alpha V$ and $\bar{E} \rightarrow \alpha \bar{E}$
 $S \rightarrow \alpha S!$ S is not extensive!

$S_T \neq \sum_i S_i$ Gibbs paradox.

Mixing 2 gases:



N_1, V_1 N_2, V_2, T
 T

Consider mixing 2
 ideal gases at the
 same T .

Before the mix:

$$S_i = N_i k \ln V_i + \frac{3}{2} N_i k \left\{ 1 + \ln \frac{2\pi m_i k T}{h^2} \right\}$$

After: $i=1,2$ $V = V_1 + V_2$

$$\textcircled{*} S_{\text{total}} = \sum_{i=1}^2 \left[N_i k \ln V + \frac{3}{2} N_i k \left\{ 1 + \ln \frac{2\pi m_i k T}{h^2} \right\} \right]$$

$$\Delta S = S_T - \sum_{i=1}^2 S_i = k \left[N_1 \ln \frac{V_1 + V_2}{V_1} + N_2 \ln \frac{V_1 + V_2}{V_2} \right]$$

$\Delta S > 0$ irreversible process.

If $\frac{N_1}{V_1} = \frac{N_2}{V_2}$ then

$$\frac{V_1 + V_2}{V_1} = \frac{\cancel{V_1} \frac{N_1}{\cancel{V_1}} + \cancel{V_2} \frac{N_2}{\cancel{V_2}}}{\cancel{V_1} \frac{N_1}{\cancel{V_1}}} = \frac{N_1 + N_2}{N_1}$$

$$\frac{V_1 + V_2}{V_2} = \frac{N_1 + N_2}{N_2}$$

Then we obtain ΔS^* (change of S for equal n gas)

$$\Delta S^* = k \left[N_1 \ln \left(\frac{N_1 + N_2}{N_1} \right) + N_2 \ln \left(\frac{N_1 + N_2}{N_2} \right) \right] > 0$$

If the 2 gases are the same $m_1 = m_2$

and

$$S_T = N k \ln V + \frac{3}{2} N k \left\{ 1 + \ln \left(\frac{2\pi m k T}{h^2} \right) \right\}$$

$$N = N_1 + N_2 \quad \text{and} \quad V = v_1 + v_2$$

$$\text{If } \frac{N_1}{v_1} = \frac{N_2}{v_2}$$

$$\Delta S^* > 0 !$$

This is wrong!

Notice that now

$$\begin{aligned}
 \Delta S^* &= k [N_1 \ln(N_1 + N_2) - N_1 \ln N_1 + \\
 &N_2 \ln(N_1 + N_2) - N_2 \ln N_2] = \\
 &= k [(N_1 + N_2) \ln(N_1 + N_2) - N_1 \ln N_1 - N_2 \ln N_2] \\
 &= k [(N_1 + N_2) \ln(N_1 + N_2) - (N_1 + N_2) - (N_1 \ln N_1 - N_1) \\
 &- (N_2 \ln N_2 - N_2)] \approx k [\ln(N_1 + N_2)! - \ln N_1! - \ln N_2!] \\
 &\neq 0
 \end{aligned}$$

Stirling: $\ln n! \approx n \ln n - n$
(large n)

if N_i
very large

Gibbs decided to "fix" the problem by subtracting $k \ln N! \approx k(N \ln N - N)$ from the entropy.

This corresponds to fixing Ω by dividing it by $N!$

$$\Omega \rightarrow \frac{\Omega}{N!}$$

$N!$ takes into account the overcounting of available states because the particles are not distinguishable (assuming that no state has more than 1 particle in it; OK for ideal gas).

Then

$$S = S - (Nk \ln N - Nk) = Nk \ln \left[\frac{V}{h^3} \left(\frac{4\pi m^3}{3N} \right)^{3/2} \right]$$

wrong S

$$+ \frac{3}{2} Nk - Nk \ln N + Nk =$$

$$S = Nk \ln \left[\frac{V}{N h^3} \left(\frac{4\pi m^3}{3N} \right)^{3/2} \right] + \frac{5}{2} Nk \quad \otimes$$

$$\oplus S = Nk \ln \left(\frac{V}{N} \right) + \frac{3}{2} Nk \left\{ \frac{5}{2} + \ln \left(\frac{2\pi m k T}{h^2} \right) \right\}$$

Now S is extensive since

for $V \rightarrow \alpha V$ and $N \rightarrow \alpha N \Rightarrow S \rightarrow \alpha S$.

also $\Delta S^* = 0$ for equal gases

but $\Delta S^* > 0$ (same expression as before for different)

$$E = \frac{3}{2} kT$$

Now from \otimes

$$E(N, V, S) = \frac{3 h^2 N^{5/3}}{4 \pi m V^{2/3}} e^{\left(\frac{2S}{3Nk} - \frac{5}{3} \right)}$$

$$\mu = \left. \frac{\partial \bar{E}}{\partial N} \right|_{V, S} = \bar{E} \left[\frac{5}{3N} - \frac{2S}{3N^2 k} \right]$$

since $\bar{E} = \frac{3}{2} kT$ and $P = \frac{2}{3} \frac{\bar{E}}{V} = \frac{NkT}{V}$

$$\begin{aligned} \mu &= \frac{1}{N} \left[\frac{5}{3} \bar{E} - \frac{2 \bar{E} S}{3Nk} \right] = \frac{1}{N} \left[\frac{5}{3} \bar{E} - \frac{2}{3Nk} \frac{3}{2} kT S \right] \\ &= \frac{1}{N} \left[\bar{E} + \underbrace{\frac{2}{3} \bar{E}}_{NkT = PV} - TS \right] = \frac{1}{N} [E + PV - TS] = \frac{G}{N} \end{aligned}$$

Replacing E and S in terms of N, V and T
we get

$$\mu = kT \ln \left\{ \frac{N}{V} \left(\frac{h^2}{2\pi m kT} \right)^{3/2} \right\}$$

and $F = E - TS = G - PV = \frac{M}{N} - NkT =$
 $= NkT \left\{ \ln \left[\frac{N}{V} \left(\frac{h^2}{2\pi m kT} \right)^{3/2} \right] \right\}.$

Correct enumeration of states:

Gibbs proposed $\Omega \rightarrow \frac{\Omega}{N!}$

it works if we consider N indistinguishable particles in systems in which each level has one particle at most (correct for ideal gas).

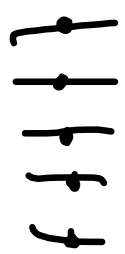
But if we consider a distribution $\{n_i\}$ where $n_i > 1$ for some i 's there are fewer configurations so Gibbs recipe is too naïve.

If $n_i > 1$ we have $\frac{N!}{n_1! n_2! \dots}$




$$\{m_i\} \quad \sum_i m_i = N$$

$$\sum_i \epsilon_i m_i \bar{G}$$

Example: $N=5$


 $N! = 120$
 configurations
 that are not
 different so
 we need to
 correct

if

n_3		1	$5 \binom{4}{2}$
n_2		2	$= \frac{5 \times 4 \times 3}{2 \times 2}$
n_1		2	$= 30$

$$\frac{N!}{n_1! n_2!} = \frac{5!}{2! 2!} = 30$$

$$\Omega \{n_i\} = \frac{1}{n_1! n_2! \dots} \quad \text{needs to be}$$

considered if n is large or T is very low.
(for non ideal gas)

Quantum mechanics will take care of
this.

Ensemble Theory

Macrostate \longrightarrow very large $\#$ of
 N, V, \bar{E} equally likely
microstates compatible
with N, V, \bar{E}

Thermodynamic properties are obtained from
a time average of the macrostate.

Ensemble picture: consider N systems representing
each of the N possible microstates.
This group of systems forms an
ensemble.

The average at a fixed time t over the members of the ensemble provides the thermodynamical properties.

Classical System: Phase space.

The position \bar{q}_i and momenta \bar{p}_i at a given time t of each of the N particles that form a classical system provide the microstate of the system.

Then phase space has $3N + 3N = 6N$
 dimensions ($N \sim 10^{23}$)

We will label:

$$\{q_i\} = q_1 \dots q_{3N}$$

$$\{p_i\} = p_1 \dots p_{3N}$$

then (q_i, p_i) $i = 1, \dots, 3N$ defines a point
 (microstate) in
 phase space.

p_i and q_i are $p_i(t)$ and $q_i(t)$.

The time dependence is given by:

$$\dot{q}_i = \frac{\partial H(q_i, p_i)}{\partial p_i} \quad (1)$$

$$\dot{p}_i = - \frac{\partial H(q_i, p_i)}{\partial q_i} \quad (2)$$

$H(q_i, p_i)$ is the system's Hamiltonian.

Each (q_i, p_i) evolves as a function of t describing a trajectory in phase space.

If E in the system is fixed we know that $H(q_i, p_i) = E$ for all t .

The trajectory of (q_i, p_i) will define a hypersurface in phase space (not all microstates will be possible).

If E is given within a range

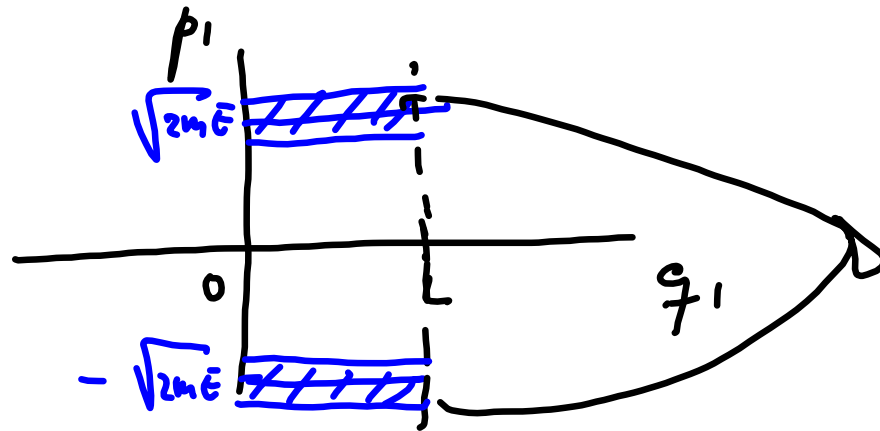
$E - \frac{1}{2} \Delta \leq E \leq E + \frac{1}{2} \Delta$ the accessible points define a shell of width Δ .

Example: $N=1$ 1D free particle.

$$E = \frac{p_1^2}{2m}$$

box of length L .

$$p_1 = \pm \sqrt{2mE}$$



hypershell of
accessible
states.