

9/29

Last time:

Grand canonical ensemble.

$$\tilde{Z} = \sum_{N_r=0}^{\infty} \lambda^{N_r} Z_{N_r}(V, T)$$

$$U(\lambda, V, T) = kT^2 \left[\frac{\partial}{\partial T} \ln \tilde{Z}(\lambda, V, T) \right]$$

$$\ln \tilde{Z}$$

$$F = -kT \ln \tilde{Z} / \lambda^N$$

Then

$$S = \frac{U - F}{T} = \frac{kT^2}{T} \left[\frac{\partial}{\partial T} \ln \tilde{Z}(\lambda, V, T) \right] + \frac{NkT}{T} \ln \lambda + kq$$

Then

$$S = kT \left. \frac{\partial \ln Z}{\partial T} \right|_{z, v} - Nk \ln z + k \ln Z.$$

Examples:

1) Classical ideal gas:

We found that the canonical Z for the ideal gas was given by

$$Z_N(V, T) = \frac{Z_1^N(V, T)}{N!}$$

do take into account the fact that the particles are indistinguishable!

$$Z_1(V, T) = V f(T) \quad \textcircled{1}$$

↳ the molecule can be anywhere inside V .

Then

$$Z(N, V, T) = \sum_{N_r=0}^{\infty} \frac{Z_1^{N_r}}{N_r!} =$$

$$\stackrel{\textcircled{1}}{=} \sum_{N_r=0}^{\infty} \frac{(V f(T))^{N_r}}{N_r!} = e^{V f(T)}$$

$$e^x = \sum_{j=0}^{\infty} \frac{x^j}{j!}$$

Then

$$\Omega(\beta, \nu, T) = -kT \ln \mathcal{Z} = -kT \int V f(T) = -PV$$

↳
Landau potential

also

$$g(\beta, \nu, T) = \ln \mathcal{Z} = \int V f(T) = \frac{PV}{kT}$$

$$\therefore P = \frac{kT}{V} g = \frac{kT}{V} \int V f(T) = kT \int f(T) \quad (2)$$

Also (see previous lecture)

$$N = \int \frac{\partial}{\partial \beta} g(\beta, \nu, T) \Big|_{\nu, T} = \int V f(T) \quad (3)$$

$$U = kT^2 \left[\frac{\partial \ln \zeta(z, V, T)}{\partial T} \right]_{z, V} = kT^2 \partial V \frac{\partial \ln \zeta}{\partial T} \quad (4)$$

$$F = NkT \ln \zeta - kT \underbrace{\partial V f(T)}_f \quad (5)$$

$$\begin{aligned} S = \frac{U - F}{T} &= kT \partial V f' - Nk \ln \zeta + k \partial V f = \\ &= -Nk \ln \zeta + \partial V k (T f' + f) \quad (6) \end{aligned}$$

From (2) and (3) we obtain:

$$\zeta = \frac{P}{kTf} \quad \zeta = \frac{N}{Vf} \Rightarrow \frac{P}{kT} = \frac{N}{V}$$

or $\boxed{PV = NkT}$

From (3) and (4) we get

$$\frac{N}{Vf} = \beta = \frac{U}{kT^2 V f'} \Rightarrow U = NkT^2 \frac{f'}{f}$$

If $f(T) \propto T^m$ then

$$U = NkT^2 \mu \frac{T^{m-1}}{T^m} = \frac{NkT^2 \mu}{T} = \boxed{NkT\mu}$$

If the gas is non-relativistic $\mu = 3/2$

$$U = \frac{3}{2} NkT$$

If gas is extremely relativistic $\mu = 3$
and $U = 3NkT$.

Fluctuations of \bar{N} and \bar{E} in the grand-canonical

$$\bar{N} = \frac{\sum_{r,s} N_r e^{-\alpha N_r - \beta E_s}}{\sum_{r,s} e^{-\alpha N_r - \beta E_s}}$$

$$\left. \frac{\partial \bar{N}}{\partial \alpha} \right|_{\beta, E_s} = \frac{-\sum_{r,s} N_r^2 e^{-\alpha N_r - \beta E_s}}{\sum_{r,s} e^{-\alpha N_r - \beta E_s}} +$$

$$+ \frac{\sum_{r,s} N_r e^{-\alpha N_r - \beta E_s}}{\sum_{r,s} e^{-\alpha N_r - \beta E_s}} \quad (+) \quad \frac{\sum_{r,s} N_r e^{-\alpha N_r - \beta E_s}}{\sum_{r,s} e^{-\alpha N_r - \beta E_s}} =$$

$$= -\langle N_r^2 \rangle + \langle N_r \rangle^2$$

Then

$$\overline{(\Delta N)^2} = \overline{N^2} - (\overline{N})^2 = - \frac{\partial \overline{N}}{\partial \alpha} \Big|_{\beta, \epsilon_s} = kT \frac{\partial \overline{N}}{\partial \mu} \Big|_{T, V}$$

Also define $n = \frac{N}{V}$

$$\frac{\overline{(\Delta n)^2}}{(\overline{n})^2} = \frac{\overline{(\Delta N)^2}}{(\overline{N})^2} = \frac{kT}{(\overline{N})^2} \frac{\partial \overline{N}}{\partial \mu} \Big|_{T, V} = \frac{kT n^2}{V^2} \frac{\partial (V/n)}{\partial \mu} \Big|_{T, V}$$

$$= \frac{kT}{V^2} n^2 \checkmark \frac{\partial n^{-1}}{\partial \mu} \Big|_{T, V} = - \frac{kT}{V} \frac{\partial n}{\partial \mu} \Big|_{T, V} \quad \textcircled{7}$$

$\underbrace{\quad}_{-\frac{1}{n^2} \frac{\partial n}{\partial \mu}}$

$$\alpha = -\mu/kT$$

$$n = V/\overline{N} \Rightarrow \overline{N} = \frac{V}{n}$$

we need to find a relationship between μ and N :

From first law:

$$d\bar{E} = TdS - PdV + \mu dN$$

But \bar{E} is extensive

$$d(\lambda\bar{E}) = Td(\lambda S) - Pd(\lambda V) + \mu d(\lambda N)$$

$$\bar{E}d\lambda + \lambda d\bar{E} = TSd\lambda + \lambda TdS - PVd\lambda - \lambda PdV +$$

$$+ \mu N d\lambda + \lambda \mu dN = d\bar{E}$$

$$= \lambda (TdS - PdV + \mu dN) +$$

$$+ (TS - PV + \mu N) d\lambda$$

Since $E = TS - PV + \mu N$

then $dE = TdS + SdT - PdV - VdP +$
 $+ \mu dN + Nd\mu$

but since we know that $dE = TdS - PdV + \mu dN$

it means that

$$SdT - VdP + Nd\mu = 0 \quad \text{constraint}$$

then $d\mu = \frac{V}{N}dP - \frac{S}{N}dT = \nu dP - s dT$ (8)

Going back to (7):

$$\frac{\overline{(\Delta n)^2}}{(\bar{n})^2} = -\frac{kT}{V} \left. \frac{\partial N}{\partial \mu} \right|_{T, V}$$

since from (8) $d\mu = v dP - s dT$ if $dT = 0$

$$\Rightarrow d\mu = v dP \Rightarrow$$

$$\frac{\overline{(\Delta n)^2}}{(\bar{n})^2} = -\frac{kT}{V} \left. \frac{\partial N}{\partial P} \right|_{T, V} = -\frac{kT}{V} K_T = -\frac{P}{N} K_T \alpha\left(\frac{1}{V}\right)$$

K_T : isothermal
compressibility (intrinsic).

The fluctuations $\rightarrow 0$ when $N \rightarrow \infty$.

$$\overline{(\Delta \bar{E})^2} = \langle \bar{E}^2 \rangle - \langle \bar{E} \rangle^2 = - \frac{\partial \bar{E}}{\partial \beta} \Big|_{\beta, \nu} = k T^2 \frac{\partial U}{\partial T} \Big|_{\beta, \nu} \quad (9)$$

$U \equiv \bar{E}$

but

$$\frac{\partial U}{\partial T} \Big|_{\beta, \nu} = \frac{\partial U}{\partial T} \Big|_{N, \nu} + \frac{\partial U}{\partial N} \Big|_{T, \nu} \frac{\partial N}{\partial T} \Big|_{\beta, \nu} \quad (10)$$

Since

$$N = - \frac{\partial \ln \tilde{Z}}{\partial \alpha} \Big|_{\beta, \nu} \quad \text{and} \quad U = - \frac{\partial \ln \tilde{Z}}{\partial \beta} \Big|_{\alpha, \nu}$$

$$\frac{\partial N}{\partial \beta} \Big|_{\alpha, \nu} = \frac{\partial U}{\partial \alpha} \Big|_{\beta, \nu}$$

$$\frac{\partial N}{\partial \beta} \Big|_{\alpha, \nu} = \frac{\partial U}{\partial \alpha} \Big|_{\beta, \nu}$$

since $\beta = \frac{1}{kT}$ and $\alpha = \frac{-\mu}{kT}$

$$\frac{\partial N}{\partial T} \Big|_{\alpha, \nu} \frac{\partial T}{\partial \beta} \Big|_{\alpha, \nu} = \frac{\partial U}{\partial \mu} \Big|_{T, \nu} \frac{\partial \mu}{\partial \alpha} \Big|_{T, \nu}$$

$$-kT^2 \frac{\partial N}{\partial T} \Big|_{\alpha, \nu} = -kT \frac{\partial U}{\partial \mu} \Big|_{T, \nu}$$

$$\therefore \frac{\partial N}{\partial T} \Big|_{\beta, \nu} = \frac{1}{T} \frac{\partial U}{\partial \mu} \Big|_{T, \nu} \quad (11)$$

Plugging (11) in (10) and (10) in (9):

$$\overline{(\Delta E)^2} = kT^2 \left[\underbrace{\frac{\partial U}{\partial T}}_{C_V} \Big|_{N,V} + \frac{\partial U}{\partial N} \Big|_{T,V} \frac{1}{T} \frac{\partial U}{\partial \mu} \Big|_{T,V} \right]$$

$$= \underbrace{kT^2 C_V}_{\langle (\Delta E)^2 \rangle_{\text{canonical}}} + kT \frac{\partial U}{\partial N} \Big|_{T,V} \underbrace{\frac{\partial U}{\partial \mu} \Big|_{T,V}}_{\frac{\partial U}{\partial N} \Big|_{T,V} \frac{\partial N}{\partial \mu} \Big|_{T,V}} =$$

$$= \langle (\Delta E)^2 \rangle_{\text{canonical}} + \left(\frac{\partial U}{\partial N} \Big|_{T,V} \right)^2 \underbrace{\overline{(\Delta N)^2}}_{kT \frac{\partial N}{\partial \mu} \Big|_{T,V}}$$

Vanish when $N \rightarrow \infty$.

Quantum Statistics

Density Matrix:

$N \gg 1$ ensemble members.

\hat{H} hamiltonian operator

At any time t $\psi(\vec{r}_i, t)$ is the wave function for the ensemble.

Each ensemble member has its own wave function $\psi^k(r_i, t)$ $k = 1, \dots, N$

The ψ^k 's obey Schrödinger's equation:

$$\hat{H} \psi^k(t) = i\hbar \dot{\psi}^k(t) \quad (1)$$

$$\psi^k(t) = \sum_n a_n^k(t) \phi_n \quad (2)$$

ϕ_n : functions in an orthonormal basis.

time independent

since $\int \phi_m^* \phi_n d\tau = \delta_{m,n}$

$$\int \phi_m^* \psi^k(t) d\tau = \sum_n a_n^k(t) \underbrace{\int \phi_m^* \phi_n d\tau}_{\delta_{m,n}}$$

$$= a_m^k(t)$$

or

$$a_n^k(t) = \int \phi_n^* \psi^k(t) d\tau \quad (3)$$

Let's find the time evolution of $a_n^k(t)$:

$$i\hbar \dot{a}_n^k(t) \stackrel{(3)}{=} i\hbar \int \phi_n^* \dot{\psi}^k(t) d\mathcal{V} \stackrel{(1)}{=} \quad (1)$$

$$\stackrel{(1)}{=} \frac{i\hbar}{i\hbar} \int \phi_n^* \hat{H} \psi^k(t) d\mathcal{V} \stackrel{(2)}{=} \quad (2)$$

$$\stackrel{(2)}{=} \int \phi_n^* \hat{H} \sum_m a_m^k(t) \phi_m d\mathcal{V} =$$

$$= \sum_m a_m^k(t) \underbrace{\int \phi_n^* \hat{H} \phi_m d\mathcal{V}} =$$


$$= \sum_m H_{nm} a_m^k(t) \quad H_{nm} \quad (4)$$

From (2) we see that $|a_m^k(t)|^2$ is the probability of finding the system k in eigenstate m at time t . Then

$$\sum_m |a_m^k(t)|^2 = 1 \quad \forall k$$

Define $\hat{\rho}(t)$ such that

$$\rho_{mm}(t) = \frac{1}{N} \sum_{k=1}^N \left\{ a_m^k(t) (a_n^k(t))^* \right\} \quad (5)$$



ensemble average

Notice that $\rho_{nn}(t)$ is $\langle a_n^k a_n^{k*} \rangle =$
 $= \langle |a_n|^2 \rangle \rightarrow$ ensemble average.

Then $\rho_{nn}(t)$ is the probability that a system chosen at random from the ensemble is in state n at time t \therefore

$$\sum_n \rho_{nn} = 1$$

because the system has to be in some of the n states.

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$$\sum_m \rho_{mm} = \sum_m \frac{1}{N} \sum_{k=1}^N \{ a_m^k(t) a_m^{k*}(t) \} =$$

$$= \frac{1}{N} \sum_{k=1}^N \underbrace{\sum_m |a_m^k(t)|^2}_1 = \frac{N}{N} = 1$$

Let's find the time dependence of $\rho_{mn}(t)$:

$$i\hbar \dot{\rho}_{mn}(t) \stackrel{\textcircled{5}}{=} i\hbar \frac{1}{N} \sum_{k=1}^N \{ \dot{a}_m^k(t) a_n^{k*}(t) +$$

$$+ a_m^k(t) \dot{a}_n^{k*}(t) \} \stackrel{\textcircled{4}}{=} \frac{1}{N} \sum_{k=1}^N \left\{ \sum_i H_{mj} a_j^k(t) a_n^{k*}(t) + \right. \\ \left. - a_m^k(t) \sum_j H_{nj}^* a_j^{k*}(t) \right\} =$$

$$= \sum_j (H_{mj} \rho_{jn} - \underbrace{H_{nj}^*}_{H_{jn} \text{ (Hermitian)}} \rho_{mj}) =$$

using (5)

$$= \sum_j (H_{mj} \rho_{jn} - \rho_{mj} H_{jn}) =$$

$$= (\hat{H} \hat{\rho} - \hat{\rho} \hat{H})_{mn}$$

$$\therefore i \hbar \dot{\hat{\rho}} = [\hat{H}, \hat{\rho}]. \quad \text{Quantum equivalent of Liouville's theorem.}$$

$$\left[\frac{\partial \rho}{\partial t} = -[\rho, H] = - \sum_{i=1}^{3N} \left(\frac{\partial \rho}{\partial q_i} \frac{\partial H}{\partial p_i} - \frac{\partial \rho}{\partial p_i} \frac{\partial H}{\partial q_i} \right) \right].$$