

System in canonical ensemble.

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N : systems sharing energy \mathcal{E} .

E_r : $r = 0, 1, 2, \dots$ possible energies of each system.

n_r : # of systems with energy E_r .

$\{n_r\}$: distribution.

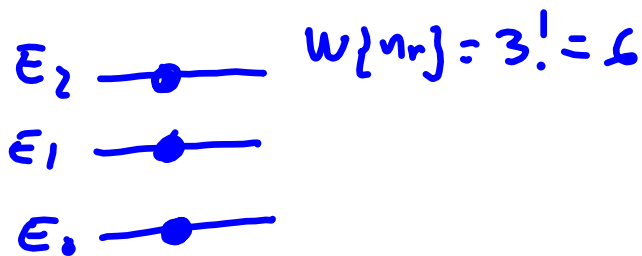
Constraints:

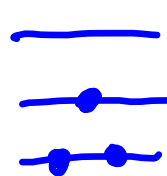
$$\textcircled{1} \left\{ \begin{array}{l} \sum_r n_r = N \\ \sum_r n_r E_r = \mathcal{E} = N U \end{array} \right. \quad \begin{array}{l} \text{average} \\ \text{energy} \\ U = \frac{\mathcal{E}}{N} \end{array}$$

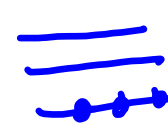
Notice that $\{n_r\}$ have a degeneracy because we can change which member of the ensemble has each energy E_r . Then the degeneracy of $\{n_r\}$ is given by:

$$W\{n_r\} = \frac{N!}{n_0! n_1! n_2! \dots}$$

Ex: $N = 3$




 $W\{n_r\} = 3 = \frac{3!}{1! 2! 0!} = \frac{6}{2} = 3$


 $W\{n_r\} = 1 = \frac{3!}{3! 0! 0!} = 1$

Now all the states of the ensemble compatible with ① are equally likely. Then the most probable $\{n_r\}$ will be the one with the largest $W\{n_r\}$.

Ex:

$$N=2 \quad \Sigma = 2\varepsilon$$

$$\text{---} \quad \varepsilon_2 = 2\varepsilon$$

$$\text{---} \quad \varepsilon_1 = \varepsilon$$

$$\text{---} \quad \varepsilon_0 = 0$$

$$\text{---} \bullet \text{---} \quad W\{n_r\} = 2$$

$$\text{---}$$

$$\text{---} \bullet \text{---}$$

higher W

will be preferred

$$\text{---}$$

$$\text{---} \bullet \bullet \text{---}$$

$$\text{---}$$

$$W\{n_r\} = 1$$

We need to find $\{n_r^*\}$ with maximum W .

We will see that W is very sharp so that if $\{n_r\} \neq \{n_r^*\}$ the probability of the distribution will be negligible.

We will find \rightarrow sum over $\{n_r\}$ schrittweise ①

$$\langle n_r \rangle = \frac{\sum'_{\{n_r\}} n_r W\{n_r\}}{\sum'_{\{n_r\}} W\{n_r\}}$$

We will find n_r^* and $\langle n_r \rangle$ and we will see that $n_r^* = \langle n_r \rangle$ when $N \rightarrow \infty$.

1) Method of most probable values:

$$\ln W = \ln N! - \sum_r \ln n_r!$$

Since N and n_r are very large we can use Stirling's approx: $\ln N! \approx N \ln N - N$

$$\begin{aligned} \therefore \ln W &= N \ln N - N - \sum_r (n_r \ln n_r - n_r) = \\ &= N \ln N - \cancel{N} - \sum_r n_r \ln n_r + \underbrace{\sum_r n_r}_{\cancel{N}} = \\ &= N \ln N - \sum_r n_r \ln n_r \end{aligned}$$

Let's find $\{n_r^*\}$ that satisfies $\delta W\{n_r^*\} = 0$

$$\begin{aligned}\delta(\ln W) &= - \sum_r \left(\ln n_r + \frac{n_r}{n_r} \right) \delta n_r = \\ &= - \sum_r (\ln n_r + 1) \delta n_r = 0\end{aligned}$$

but the sum over r has constraints and this means that:

$$\left\{ \begin{array}{l} \sum_r \delta n_r = 0 \\ \sum_r E_r \delta n_r = 0 \end{array} \right. \text{ constraints.}$$

We can include the constraints via Lagrange multipliers then

$$\delta \ln W = 0 = \sum_r \left[-(\ln m_r + 1) - \alpha - \beta \epsilon_r \right] \delta n_r$$

→ unrestricted sum

Since δn_r are now arbitrary we have that

$$-(\ln m_r + 1) - \alpha - \beta \epsilon_r = 0$$

$$\therefore \ln m_r^* = -\beta \epsilon_r - \alpha - 1$$

$$\boxed{m_r^* = C e^{-\beta \epsilon_r}} \quad (+)$$

We can use the constraints ① to find c
and β :

$$\mathcal{N} = \sum_r n_r = \sum_r n_r^* = c \sum_r e^{-\beta \bar{\epsilon}_r}$$

$$\therefore c = \frac{\mathcal{N}}{\sum_r e^{-\beta \bar{\epsilon}_r}}$$

Then $\frac{n_r^*}{\mathcal{N}} = \frac{e^{-\beta \bar{\epsilon}_r}}{\sum_r e^{-\beta \bar{\epsilon}_r}}$ ++

Also

$$N U = \mathcal{E} = \sum_r \mathcal{E}_r n_r = \sum_r \mathcal{E}_r n_r^{\pm} =$$

$$= N \frac{\sum_r \mathcal{E}_r e^{-\beta \mathcal{E}_r}}{\sum_r e^{-\beta \mathcal{E}_r}} \Rightarrow$$

$$U = \frac{\mathcal{E}}{N} = \frac{\sum_r \mathcal{E}_r e^{-\beta \mathcal{E}_r}}{\sum_r e^{-\beta \mathcal{E}_r}}$$

If you know U from the above equations you can obtain β .

Method of Mean Values:

$$\langle n_r \rangle = \frac{\sum'_{\{n_r\}} n_r W\{n_r\}}{\sum'_{\{n_r\}} W\{n_r\}}$$

Trick:

Define

$$\tilde{W}\{n_r\} = \frac{N! \omega_0^{n_0} \omega_1^{n_1} \omega_2^{n_2} \dots}{n_0! n_1! n_2! \dots}$$

Notice that $\tilde{W}\{n_r\} |_{\forall \omega_r=1} \equiv W\{n_r\}$

Define:

$$\Gamma(N, \omega) = \sum'_{\{n_r\}} \tilde{W}\{n_r\}$$

You can see that

$$\begin{aligned} \textcircled{*} \langle n_r \rangle &= \omega_r \frac{\partial}{\partial \omega_r} (\ln \Gamma) \Big|_{\forall \omega_r=1} = \\ &= \omega_r \frac{\partial}{\partial \omega_r} \ln \sum'_{\{n_r\}} \frac{N! \omega_0^{n_0} \omega_1^{n_1} \dots}{n_0! n_1! \dots} \Big|_{\forall \omega_r=1} = \\ &= \omega_r \frac{\sum'_{\{n_r\}} n_r \omega_r^{n_r-1} \tilde{W}'\{n_r\}}{\sum'_{\{n_r\}} \tilde{W}\{n_r\}} \Big|_{\forall \omega_r=1} = \frac{\sum'_{\{n_r\}} n_r W\{n_r\}}{\sum'_{\{n_r\}} W\{n_r\}} \Big|_{\forall \omega_r=1} \end{aligned}$$

The problem to evaluate Γ or W is dealing with the constraint on the sum over $\{n_r\}$:

Construct

$$G(N, z) = \sum_{U=0}^{\infty} \Gamma(N, U) z^{NU} =$$

the
generating function
power series in z
with $\Gamma(N, U)$ as the
coefficients.

$$= \sum_{U=0}^{\infty} \left[\sum_{\{n_r\}} \frac{N!}{n_0! n_1! \dots} (w_0 z^{\epsilon_0})^{n_0} (w_1 z^{\epsilon_1})^{n_1} \dots \right]$$

Using $\sum_i m_i \epsilon_i = N U$

$$G(N, \beta) = (\omega_0 z_0^{\epsilon_0} + \omega_1 z_1^{\epsilon_1} + \dots)^N = f(z)$$

Now read in the book (and ask me questions if you have them) the derivation that leads to

$$\frac{1}{N} \ln \Omega(N, U) = \ln \left\{ \sum_r \omega_r e^{-\beta \epsilon_r} \right\} + \beta U$$

Replacing $\ln \Omega(N, U)$ in the expression for $\langle n_r \rangle$ (*):

$$\begin{aligned}
 \langle n_r \rangle &= \omega_r \left. \frac{\partial \ln \mathcal{N}}{\partial \omega_r} \right|_{\forall \omega_r=1} = \\
 &= \omega_r \left. \frac{\partial}{\partial \omega_r} \left[\mathcal{N} \ln \left\{ \sum_r \omega_r e^{-\beta \epsilon_r} \right\} + \mathcal{N} \beta U \right] \right|_{\forall \omega_r=1} = \\
 &= \omega_r \left[\frac{\mathcal{N} e^{-\beta \epsilon_r}}{\sum_r \omega_r e^{-\beta \epsilon_r}} + \left\{ \frac{-\mathcal{N} \sum_r \omega_r \epsilon_r e^{-\beta \epsilon_r}}{\sum_r \omega_r e^{-\beta \epsilon_r}} + \mathcal{N} U \right\} \frac{\partial \beta}{\partial \omega_r} \right]_{\forall \omega_r=1} = \\
 &\quad \underbrace{\mathcal{N} U}_{\text{for } \omega_r=1} \\
 &= \frac{\mathcal{N} e^{-\beta \epsilon_r}}{\sum_r e^{-\beta \epsilon_r}} \Rightarrow \frac{\langle n_r \rangle}{\mathcal{N}} = \frac{e^{-\beta \epsilon_r}}{\sum_r e^{-\beta \epsilon_r}}
 \end{aligned}$$

We see that $\frac{\langle n_r \rangle}{N} \equiv \frac{\{n_r^+\}}{N}$ (in ++).

Let's calculate the fluctuations of $\langle n_r \rangle$:

$$\langle n_r^2 \rangle = \frac{\sum_{\{n_r\}} n_r^2 W(\{n_r\})}{\sum_{\{n_r\}} W(\{n_r\})} = \frac{1}{N} \left(n_r \frac{\partial}{\partial n_r} \right)^2 \Big|_{n_r=1}$$

homework

then

$$\begin{aligned} \langle (\Delta n_r)^2 \rangle &\equiv \langle (n_r - \langle n_r \rangle)^2 \rangle = \langle n_r^2 - 2n_r \langle n_r \rangle + \langle n_r \rangle^2 \rangle \\ &= \langle n_r^2 \rangle - 2\langle n_r \rangle^2 + \langle n_r \rangle^2 = \langle n_r^2 \rangle - \langle n_r \rangle^2 \end{aligned}$$

Then

$$\left\langle \left(\frac{\Delta n_r}{\langle n_r \rangle} \right)^2 \right\rangle = \frac{1}{\langle n_r \rangle} - \frac{1}{N} \left\{ 1 + \frac{(\bar{E}_r - U)^2}{\langle (E_r - U)^2 \rangle} \right\}$$

homework

You see that when $N \rightarrow \infty \Rightarrow \langle n_r \rangle \rightarrow \infty$
 so the fluctuations $\rightarrow 0$ and $\langle n_r \rangle \cong n_r^+$

Physical meaning of statistical quantities
in the canonical ensemble:

We found the canonical distribution

$$P_r = \frac{\langle n_r \rangle}{N} = \frac{e^{-\beta \epsilon_r}}{\sum_r e^{-\beta \epsilon_r}}$$



→ sum over all states
(as opposed to
accessible states
only as in
microcanonical)

β is obtained from

$$U = \frac{\sum_r \epsilon_r e^{-\beta \epsilon_r}}{\sum_r e^{-\beta \epsilon_r}} =$$

$$= -\frac{\partial}{\partial \beta} \ln \left\{ \sum_r e^{-\beta \epsilon_r} \right\}$$

Make contact with thermodynamics.

We know:

$$F = U - TS$$

$$dF = dU - TdS - SdT = \cancel{TdS} - PdV + \mu dN - \cancel{TdS} - SdT = -PdV - SdT + \mu dN$$

$$\therefore \boxed{S = - \frac{\partial F}{\partial T} \Big|_{V, N}} \quad P = - \frac{\partial F}{\partial V} \Big|_{T, N} \quad \mu = \frac{\partial F}{\partial N} \Big|_{V, T}$$

V, N, T are the variables that control the canonical ensemble.

We know that

$$U = F + TS = F - T \frac{\partial F}{\partial T} \Big|_{V, N} =$$

$$= -T^2 \left[\frac{\partial}{\partial T} \left(\frac{F}{T} \right)_{N, V} \right] = \left[\frac{\partial (F/T)}{\partial (1/T)} \right]_{N, V}$$

But we also found that

$$U = - \frac{\partial}{\partial \beta} \ln \left\{ \sum_r e^{-\beta \epsilon_r} \right\}$$

Comparing the 2 expressions if we identify

$$\beta = \frac{1}{kT} \text{ then } \ln \sum_r e^{-\beta \epsilon_r} = - \frac{F}{kT}$$

Then

$$F(N, V, T) = -kT \ln \sum_r e^{-\beta \epsilon_r} =$$

$$= -kT \ln Z_N(V, T)$$

$$Z_N(V, T) = \sum_r e^{-\beta \epsilon_r} = \sum_r e^{-\frac{\epsilon_r}{kT}}$$

partition function

(caution: Pathria uses Q for Z !)

Notice that all the states r appear in the sum but the weight (probability) will kill most of them.

Now if you find F from Z_U you
can find all the other properties through
derivatives of F .

$$C_V = \left. \frac{\partial U}{\partial T} \right|_{N,V} = -T \left. \frac{\partial^2 F}{\partial T^2} \right|_{N,V}, \text{ etc.}$$