

Last time:

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$$S = -k \langle \ln P_r \rangle = -k \sum_r P_r \ln P_r$$

\mathcal{F} microcanonical distribution:

$$P_r = \frac{1}{\Omega} \quad \begin{array}{l} \text{if } r \text{ is in the allowed region} \\ 0 \text{ if } r \text{ is not allowed.} \end{array}$$

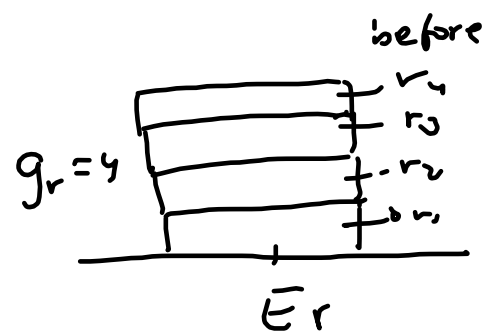
$$\begin{aligned} S &= -k \sum_{\substack{r=1 \\ \text{of} \\ \text{allowed}}}^{\Omega} \frac{1}{\Omega} \ln \left(\frac{1}{\Omega} \right) = -k \sum_{r=1}^{\Omega} \frac{1}{\Omega} (-\ln \Omega) = \\ &= k \frac{\ln \Omega}{\Omega} \underbrace{\sum_{r=1}^{\Omega} 1}_{\Omega} = k \ln \Omega \end{aligned}$$

expression of
S in microcanonical.

Partition function and density of states:

If \bar{E}_r has degeneracy g_r then

$$Z_N(V, T) = \sum_r g_r e^{-\beta \bar{E}_r}$$



now we consider the different values of \bar{E}_r

(while before many of the \bar{E}_r 's may have been identical to each other)

and

$$P_r = \frac{g_r e^{-\beta \bar{E}_r}}{\sum_r g_r e^{-\beta \bar{E}_r}}$$

For a macroscopic state $\frac{\Delta E}{E} \ll 1$, then E can be considered continuous.

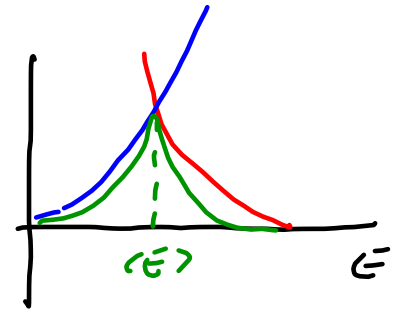
Then $P(E) dE$: probability of finding the system with energy in the interval $(E, E+dE)$.

$$P(E) dE \propto e^{-\beta E} g(E) dE$$

↓
density of states.

Normalizing $P(E) dE$ to 1:

$$P(E) dE = \frac{e^{-\beta E} g(E) dE}{\int_0^\infty e^{-\beta E} g(E) dE}$$



Now

$$Z_N(fV, T) = \int_0^{\infty} e^{-\beta \epsilon} g(\epsilon) d\bar{\epsilon}$$

Looks like
a Laplace's
transform.

and

$$\langle f \rangle = \sum_r f_r P_r = \frac{\sum_r f_r(\epsilon_r) g_r e^{-\beta \epsilon_r}}{\sum_r g_r e^{-\beta \epsilon_r}} \rightarrow$$

$$\rightarrow \frac{\int_0^{\infty} f(\epsilon) e^{-\beta \epsilon} g(\epsilon) d\bar{\epsilon}}{\int_0^{\infty} e^{-\beta \epsilon} g(\epsilon) d\bar{\epsilon}}$$

Laplace's transform:

$$F(s) = \int_0^{\infty} e^{-st} f(t) dt$$

$$t \geq 0$$

s is complex

$$s = \sigma + i\omega$$

$$Z_N(\tau) = \int_0^{\infty} e^{-\beta \bar{t}} g(\bar{t}) d\bar{t}$$

identifying

$$\beta \equiv s$$

$$\bar{t} = t$$

So $g(\bar{t})$ can be obtained if we know Z_N by applying the antilaplace's transform.

$$g(\bar{t}) = \frac{1}{2\pi i} \int_{\beta' - i\infty}^{\beta' + i\infty} e^{\beta \bar{t}} Z(\beta) d\beta \quad \text{for } \beta' > 0$$

$$g(\epsilon) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{(\beta' + i\beta'')\epsilon} z(\beta' + i\beta'') d\beta''$$

$\beta = \beta' + i\beta''$ integration in complex plane

In practice we are going to use tables of Laplace's transforms.

The canonical ensemble formalism for classical systems:

We will have to work in phase space.

$$\langle f \rangle = \frac{\int f(p, q) \rho(p, q) d^{3N}q d^{3N}p}{\int \rho(p, q) d^{3N}q d^{3N}p}$$

over all phase space

$$\int \rho(p, q) d^{3N}q d^{3N}p$$

much simpler and without the microcanonical constraints.

We said when we demonstrated Liouville's theorem

that $\rho(q, p) \propto e^{-\beta H(q, p)}$ satisfied $[H, \rho] = 0$

Liouville's.

$$\therefore \langle f \rangle = \frac{\int f(q, p) e^{-\beta H} dw}{\int e^{-\beta H} dw}$$

$$dw = d^{3N}q d^{3N}p$$

Then

$$Z_N(V, T) = \frac{1}{N!} \int e^{-\beta H(q, p)} \frac{dw}{h^{3N}}$$

of states in dw
 unit of phase space volume w_0
 do correct the overcounting if the N particles are indistinguishable (Gibb's recipe).

$$Z_N(V, T) = \frac{1}{N! h^{3N}} \int_{\text{all phase space}} e^{-\beta H(q, p)} dw$$

Ideal gas (classical, N particles, 3D):
 V, T (equilibrium).

$$H(p, q) = \sum_{i=1}^N \frac{p_i^2}{2m} \quad p_i^2 = (p_x)_i^2 + (p_y)_i^2 + (p_z)_i^2$$

$$\therefore Z_N(V, T) = \frac{1}{N! h^{3N}} \int_{\text{all space}} e^{-\frac{\beta}{2m} \sum_{i=1}^N p_i^2} \prod_{i=1}^N \pi d^3 q_i d^3 p_i =$$

$$= \frac{V^N}{N! h^{3N}} \left[\int e^{-\frac{p^2}{2mkT}} d^3 p \right]^N$$

$$= \frac{V^N}{N! h^{3N}} \left[4\pi \int_0^\infty p^2 e^{-\frac{p^2}{2mkT}} dp \right]^N \quad \underbrace{4\pi p^2 dp}_{\text{integration over } \theta \text{ and } \varphi}$$

Looking the defined integral on a table:

$$Z_N(V, T) = \frac{1}{N!} \left[\frac{V}{h^3} (2\pi m kT)^{3/2} \right]^N \quad \text{using } \ln N! \approx N \ln N - N$$

Then:

$$F(N, V, T) = -kT \ln Z_N(V, T) = -kT N \left\{ \ln \left[\frac{V}{h^3} (2\pi m kT)^{3/2} \right] - \ln N + 1 \right\} = kT N \left\{ \ln \left[\frac{N}{V} \left(\frac{h^2}{2\pi m kT} \right)^{3/2} \right] - 1 \right\}$$

It is identical to what we very painfully obtained using the microcanonical formalism.

Now we can obtain:

$$\mu = \frac{\partial F}{\partial N} \Big|_{V, T} \quad P = -\frac{\partial F}{\partial V} \Big|_{N, T} = \frac{NkT}{V}$$

and

$$S = - \left. \frac{\partial F}{\partial T} \right|_{V, N}$$

which will be the same as we obtained in microcanonical

How do we obtain the energy of the gas?

$$U = - \left. \frac{\partial \ln Z}{\partial \beta} \right|_{E_r \equiv \text{constant volume}} \equiv F + TS = \frac{3}{2} NkT$$

$$Z \propto T^{\frac{3N}{2}} \Rightarrow - \frac{\partial \ln Z}{\partial T} \frac{\partial T}{\partial \beta} = \frac{kT^2}{T^{3N/2}} \frac{3N}{2} T^{\frac{3N}{2}-1} \Rightarrow \frac{3kTN}{2}$$

$$\beta = \frac{1}{kT} \quad \frac{\partial \beta}{\partial T} = -\frac{1}{kT^2}$$

Also:

$$Z_N(V, T) = \frac{1}{N!} \left[\frac{V}{h^3} (2\pi m k T)^{3/2} \right]^N \equiv$$

$$\equiv \frac{[Z_1(V, T)]^N}{N!} \quad \textcircled{*}$$

Z_1 : single particle function

$\textcircled{*}$ valid if there are no interactions nor quantum correlations among the particles.

Knowing $Z_N(V, T)$ we can obtain $g(\epsilon)$ for the 3D ideal gas:

$$g(\epsilon) = \mathcal{L}^{-1}(Z(\beta))$$

$$Z_N = \frac{1}{N!} \left(\frac{V}{h^3}\right)^N \left(\frac{2\pi m}{\beta}\right)^{3N/2}$$

$$F(\beta) = \beta^{-3N/2}$$

$$\epsilon = t$$

$$\beta = s$$

Then

$$g(\epsilon) = \frac{1}{N!} \left(\frac{V}{h^3}\right)^N (2\pi m)^{3N/2} \frac{\sum \frac{3N}{2} - 1}{\left(\frac{3N}{2} - 1\right)!}$$

$$f(t) = \mathcal{L}^{-1}(F(s)) \left\{ \begin{array}{l} F(s) = \mathcal{L}(f(t)) \\ \frac{n!}{s^{n+1}} \\ \frac{\left(\frac{3N}{2} - 1\right)!}{\beta^{\frac{3N}{2}}} \end{array} \right.$$

Now that we found $g(\epsilon)$ you can check that

$$Z_N(V, T) = \int_0^\infty e^{-\beta \epsilon} g(\epsilon) d\epsilon = \frac{1}{N!} \left(\frac{V}{h^3} \right)^N \left(\frac{2\pi m}{\beta} \right)^{3N/2}$$

Also in 9/12 class we found that for
1 particle in a 3D box:

$$a(\epsilon) \approx \frac{2\pi V}{h^3} (2m)^{3/2} \epsilon^{1/2} \quad \text{single particle}$$

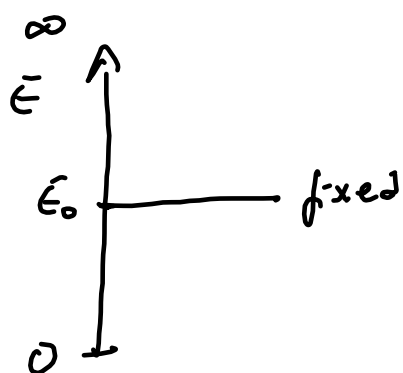
then

$$Z_1(V, T) = \frac{V}{h^3} \left(\frac{2\pi m}{\beta} \right)^{3/2} \quad \text{and} \quad Z_N = \frac{Z_1^N}{N!} .$$

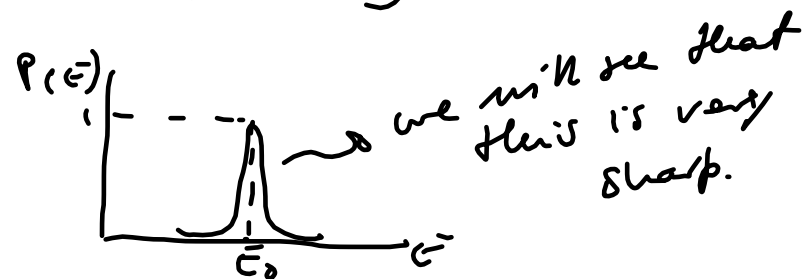
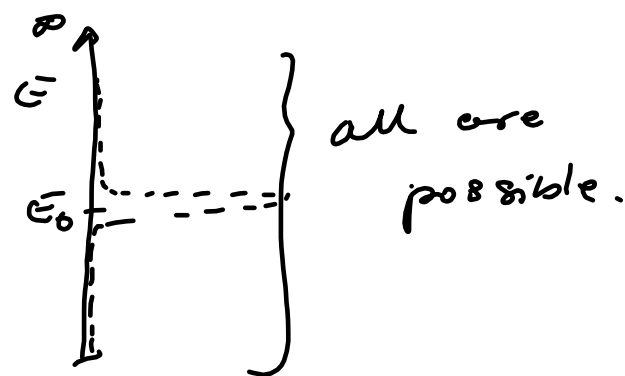
Energy fluctuations in the canonical ensemble:

Equivalence between canonical and microcanonical:

Microcanonical



Canonical



In canonical:

$$U = \langle E \rangle = \frac{\sum_r E_r e^{-\beta E_r}}{\sum_r e^{-\beta E_r}} = - \frac{\partial \ln Z}{\partial \beta}$$

$$\begin{aligned} \langle (\Delta E)^2 \rangle &= \langle (E - \langle E \rangle)^2 \rangle = \langle E^2 - 2E\langle E \rangle + \langle E \rangle^2 \rangle \\ &= \langle E^2 \rangle - 2\langle E \rangle^2 + \langle E \rangle^2 = \langle E^2 \rangle - \langle E \rangle^2 \stackrel{D/}{=} \end{aligned}$$

$$\langle E^2 \rangle = \frac{\sum_r E_r^2 e^{-\beta E_r}}{\sum_r e^{-\beta E_r}} = - \frac{\partial U}{\partial \beta}$$

$$D/ \quad \frac{\partial U}{\partial \beta} = \frac{\partial}{\partial \beta} \left[\frac{\sum_r E_r e^{-\beta E_r}}{\sum_r e^{-\beta E_r}} \right] = - \frac{\sum_r E_r^2 e^{-\beta E_r}}{\sum_r e^{-\beta E_r}} + \frac{\left(\sum_r E_r e^{-\beta E_r} \right)^2}{\left(\sum_r e^{-\beta E_r} \right)^2}$$

$$\begin{aligned} \langle (\Delta E)^2 \rangle &= -\frac{\partial U}{\partial \beta} = -\frac{\partial U}{\partial T} \frac{\partial T}{\partial \beta} = \frac{1}{k\beta^2} \frac{\partial U}{\partial T} = \\ &= \frac{(kT)^2}{k} C_V = kT^2 C_V \end{aligned}$$

Then the relative root-mean-square fluctuation is:

$$\frac{\sqrt{\langle (\Delta E)^2 \rangle}}{\langle E \rangle} = \frac{\sqrt{kT^2 C_V}}{U} \propto \frac{1}{\sqrt{N}} \xrightarrow{N \rightarrow \infty} 0$$

$U \sim N$ $C_V \sim N$

\therefore then the distribution is very sharp -
we will see that it looks like a very sharp gaussian, almost like a δ .